

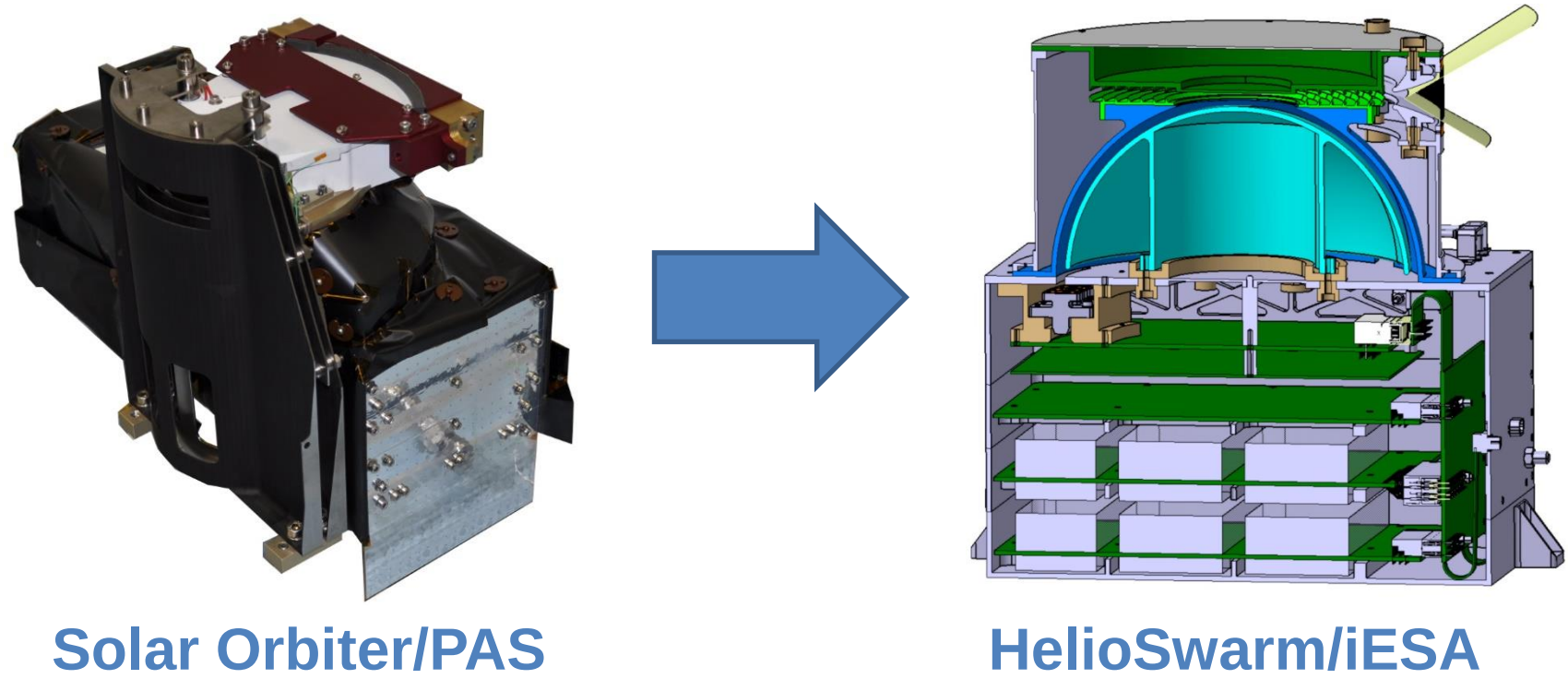


Vincent Génot (1), Benoit Lavraud (2), Naïs Fargette (1), Merlin Savard (3), Alain Barthe (4,†), Emmanuel Penou (1), King Wah Wong (1), Andrey Fedorov (1), Jack Cole (5), Kevin Hurst (5), J. Halekas (6), K. Klein (7)
(1) IRAP, CNRS, Observatoire Midi-Pyrénées, Université de Toulouse
(2) LAB, CNRS, Université de Bordeaux
(3) DPHY, ONERA, Université de Toulouse
(4) Akkodis, Toulouse
(5) Northrop Grumman
(6) University of Iowa
(7) University of Arizona

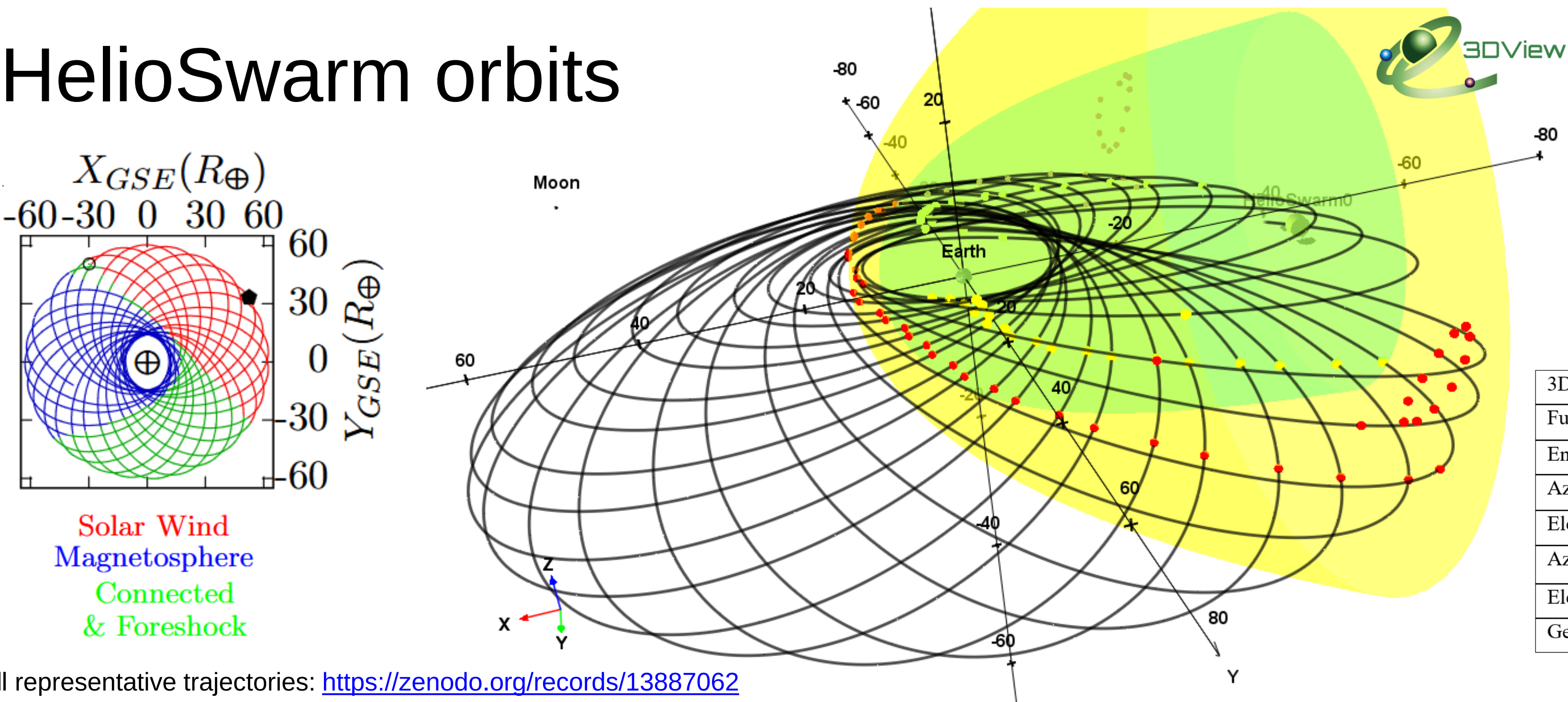
En mémoire d'Alain Barthe

We present current efforts among the HelioSwarm/iESA team to design synthetic solar wind ion data and how we use them to plan future operations. We start from Solar Orbiter/PAS data in the near-Earth environment and map them to iESA angular/energy resolutions. This synthetic dataset help us 1/ to assess the performance of the peak tracking strategy, 2/ to evaluate the scientific relevance of the different iESA modes (core, proton, alpha, full) depending on solar wind regimes, and 3/ test compression algorithms to match telemetry constrains. Visualization of distribution functions is performed with the newly developed Velocirap python library (see the related presentation).

Solar Orbiter / PAS to HelioSwarm / iESA synthetic data and operational modes



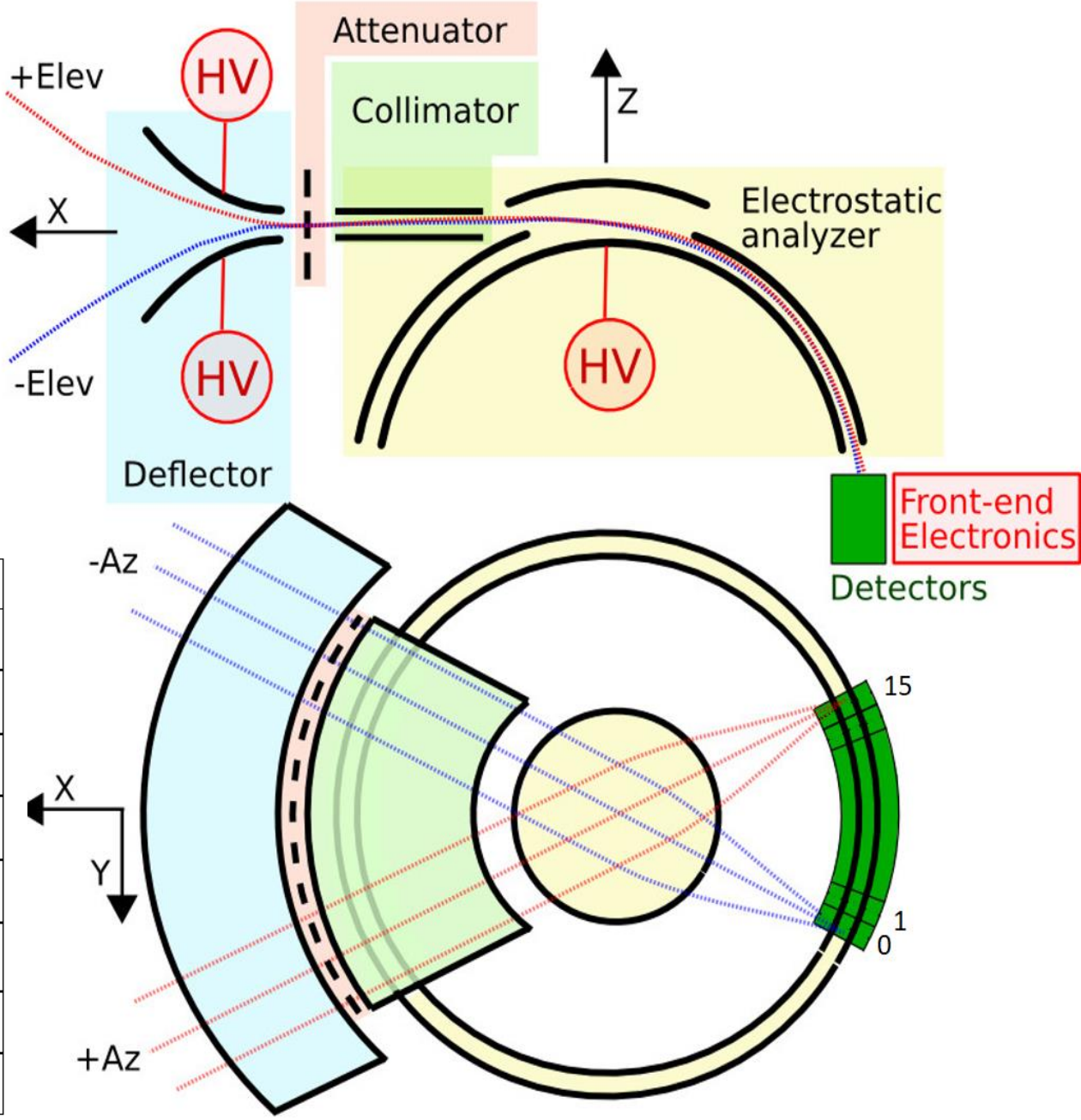
HelioSwarm orbits



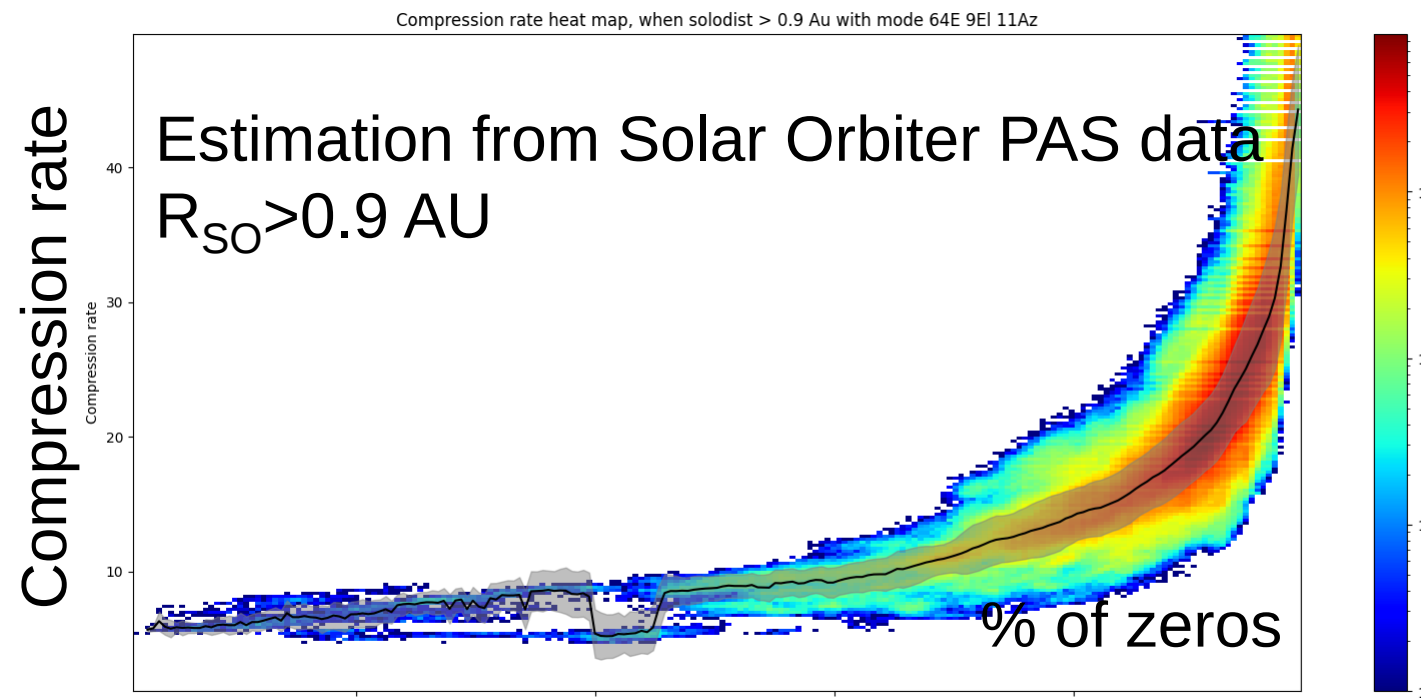
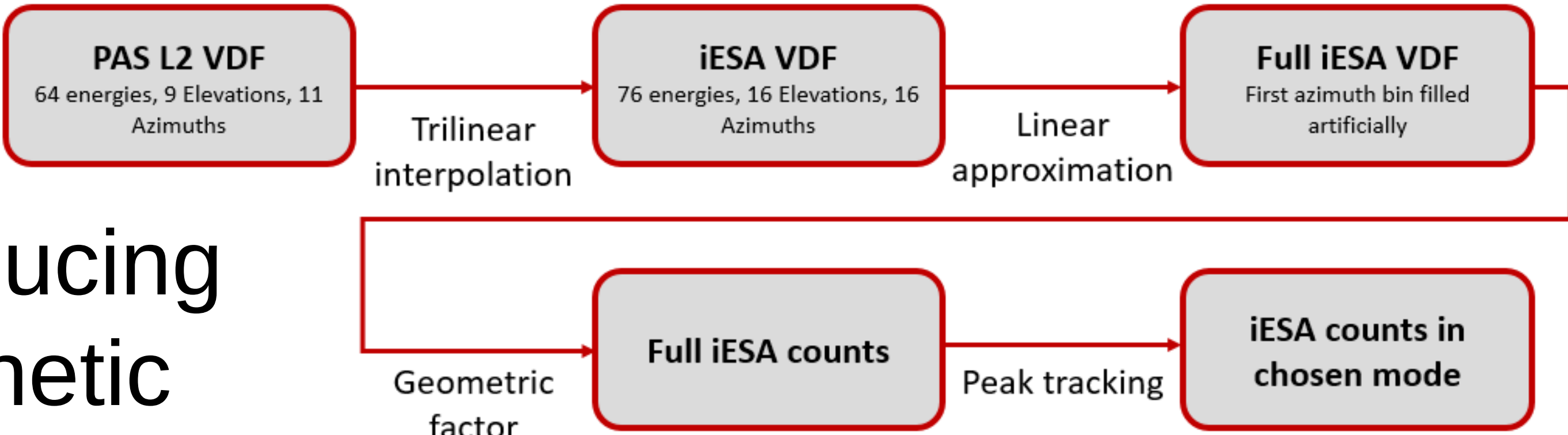
Full representative trajectories: <https://zenodo.org/records/13887062>

iESA concept

	Required	Designed
3D sampling time resolution	150ms	150 ms
Full energy range	326 eV – 6.7 keV	60 eV – 20 keV
Energy resolution $\Delta E/E$	9%	8%
Azimuth resolution	5°	3°
Elevation resolution	5°	3°
Azimuth range	$\pm 12^\circ$	$\pm 24^\circ$
Elevation range	$\pm 12^\circ$	$\pm 24^\circ$
Geometric Factor	$1.0 \cdot 10^{-6} \text{ cm}^2 \text{ sr eV/eV}$	$3.0 \cdot 10^{-6} \text{ cm}^2 \text{ sr eV/eV}$

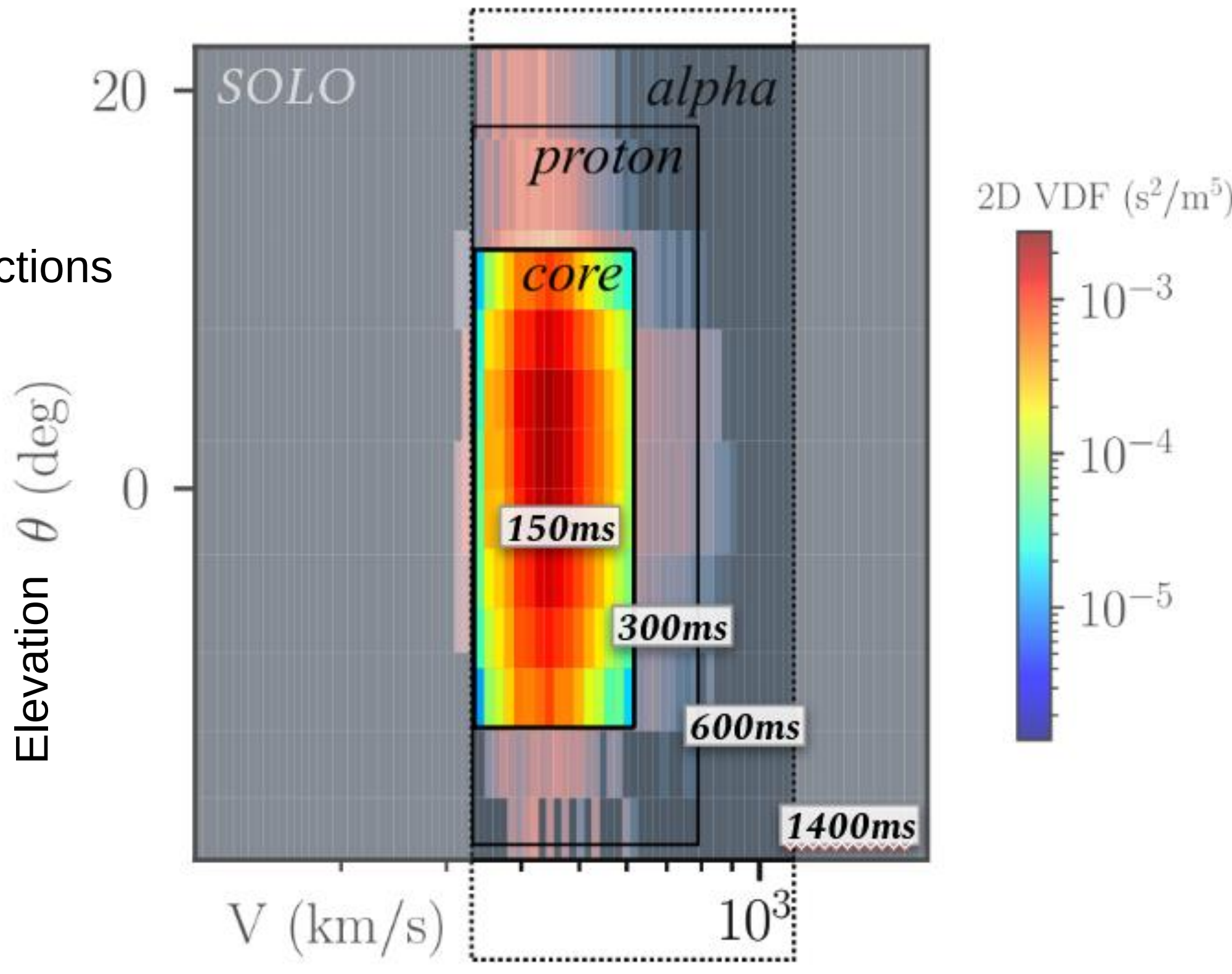


Producing synthetic iESA data



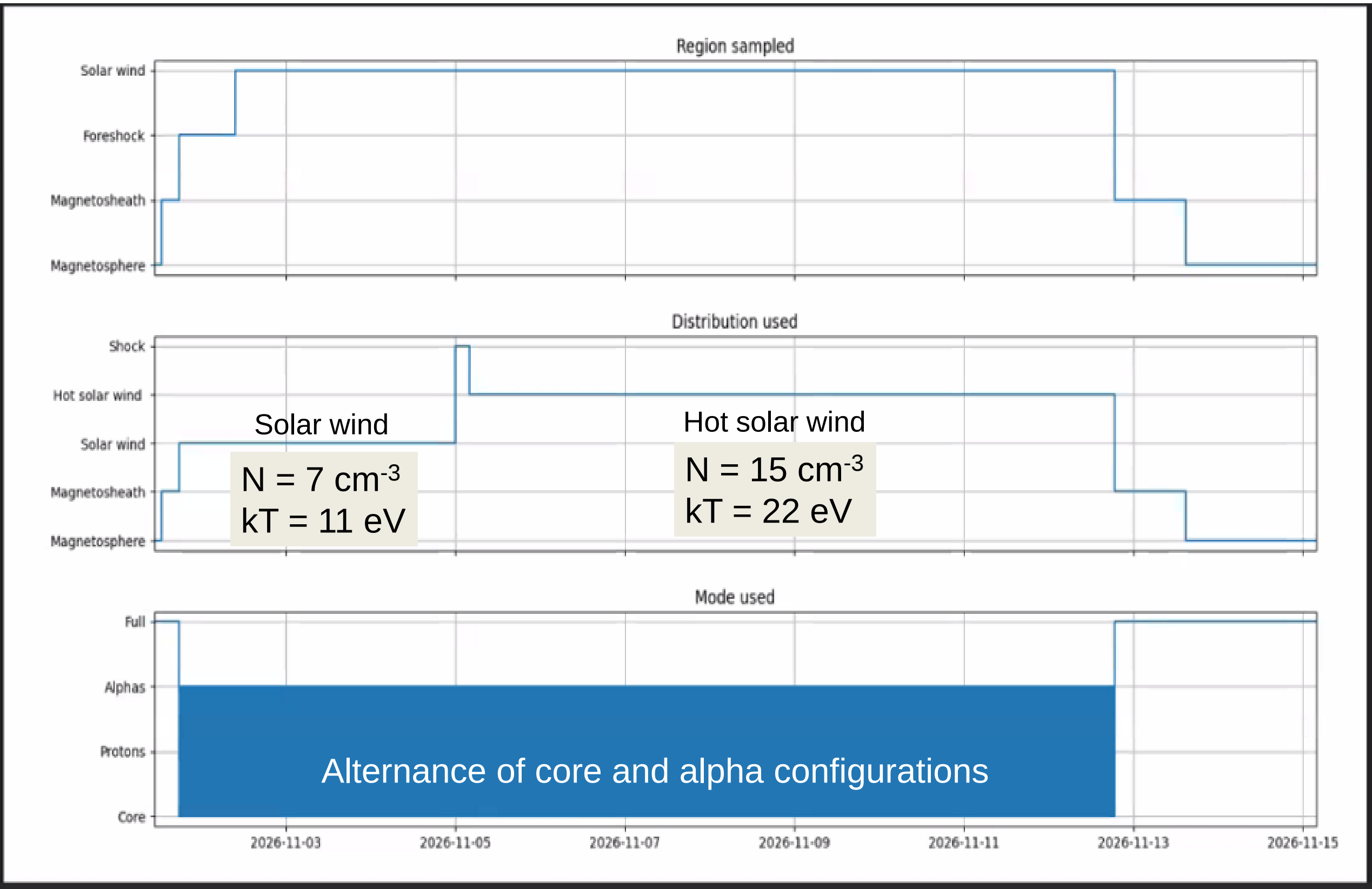
Proposed iESA configurations			
Example of possible samplings with duration in s	Sampling dimension	Sampling size at Hub PU input, Bytes	SpW rate, kBytes/s (between iESA and Hub PU)
Core configuration 0.150 s	16 energies, 8 elevations, 16 azimuths	4548	29.8 ± 0.1 (HK)
Proton configuration 0.297 s	22 energies, 12 elevations, 16 azimuths	9238	30.4 ± 0.1 (HK)
Alpha configuration 0.600 s	34 energies, 16 elevations, 16 azimuths	18890	30.8 ± 0.1 (HK)
Full configuration 1.341 s	76 energies, 16 elevations, 16 azimuths	42200	30.7 ± 0.1 (HK)

PAS and iESA distribution functions in the different configurations



Compression study over one typical HelioSwarm orbit

- For a given orbit (15 days) an iESA dataset has been generated from Solar Orbiter distribution functions in magnetosphere, magnetosheath and solar wind environments and iESA configurations
- 3 compression algorithms have been evaluated (DEFLATE, BZIP2, LZMA) on the full dataset, or by chunks of 10 MB / 10k records



10k Samples					10MB Sample Size				
Sample Count:	660				3136				
Ratio ² x to 1	Average	7.17	11.87	9.42	Average	21.12	41.50	27.30	
	Median	4.00	5.00	5.00	Median	4.00	5.00	5.00	
	Mode	4.00	5.00	5.00	Mode	4.00	5.00	5.00	
	Mean (trim top/bottom 2.5%)	5.09	7.15	6.77	Mean (trim top/bottom 2.5%)	19.14	37.10	24.83	
	Average (trim >=100)	5.36	7.49	7.13	Average (trim >=100)	8.00	12.51	10.78	
	Min	4.00	5.00	5.00	Min	4.00	5.00	5.00	
	Max	114.00	260.00	144.00	Max	114.00	248.00	144.00	
Time ⁴ Seconds	Average	2.19	1.64	25.46	Average	0.46	0.33	0.03	
	Median	2.58	1.83	27.26	Median	0.63	0.42	0.03	
	Mode	1.05	0.88	17.28	Mode	0.27	0.06	0.03	
	Mean (trim top/bottom 2.5%)	2.20	1.61	25.61	Mean (trim top/bottom 2.5%)	0.47	0.33	0.03	
	Min	0.06	0.01	0.03	Min	0.02	0.00	0.01	
	Max	3.65	4.47	40.16	Max	0.85	0.59	0.09	
Size ⁵ Bytes	Average	8,701,850.70	6,040,085.94	6,169,332.61	Average	1,831,770.99	1,272,481.29	1,309,960.24	
	Median	10,227,370.00	7,321,637.50	7,324,167.50	Median	2,582,125.00	1,849,261.00	1,866,046.00	
	Mode	10229761	3523598	7328383	Mode	92136	899673	1888642	
	Mean (trim top/bottom 2.5%)	8,815,815.85	6,128,553.33	6,244,470.03	Mean (trim top/bottom 2.5%)	1,856,933.18	1,289,025.72	1,327,270.55	
	Min	245,486.00	151,708.00	170,453.00	Min	17,605.00	9,401.00	14,201.00	
	Max	10,245,838.00	7,336,646.00	7,338,027.00	Max	2,616,753.00	1,874,668.00	1,890,500.00	

- The 3 compression algorithms provide a trade off in compression time and compressed size
- For compression, BZIP2 and LZMA provided similar performance, with the worst cases being a 5:1 compression ratio during the Hot Solar Wind sections of the example data
- When not in Hot Solar Wind, the compression was often greater than 10 to 1 for both algorithms
- When considering time, BZIP2 provides the best compression time for large files
- It is recommended that the BZIP2 algorithm is implemented in HelioSwarm FSW
- There is an existing open-source C library for BZIP2 that could be utilized
- The main risk with using BZIP2 on HelioSwarm is that it would be new development
- BZIP2 is a slower algorithm that uses more memory than DEFLATE
- BZIP2 would require more testing to ensure it does not overwhelm processing time or memory capacity

