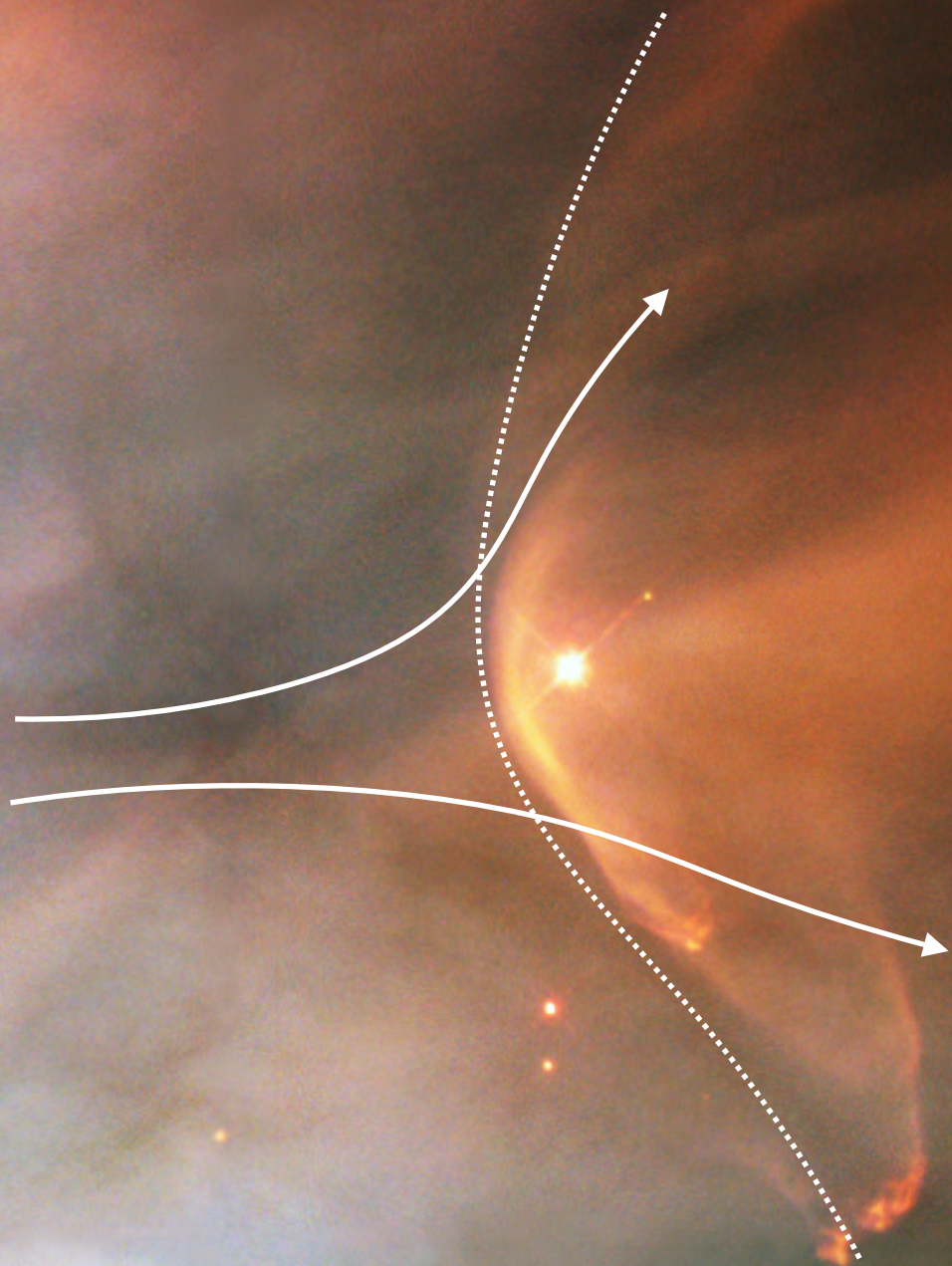


Parallel Hybrid Particle-In-Cell code with Adaptive Mesh Refinement



N. Aunai

Philip Deegan

Thibault Payet

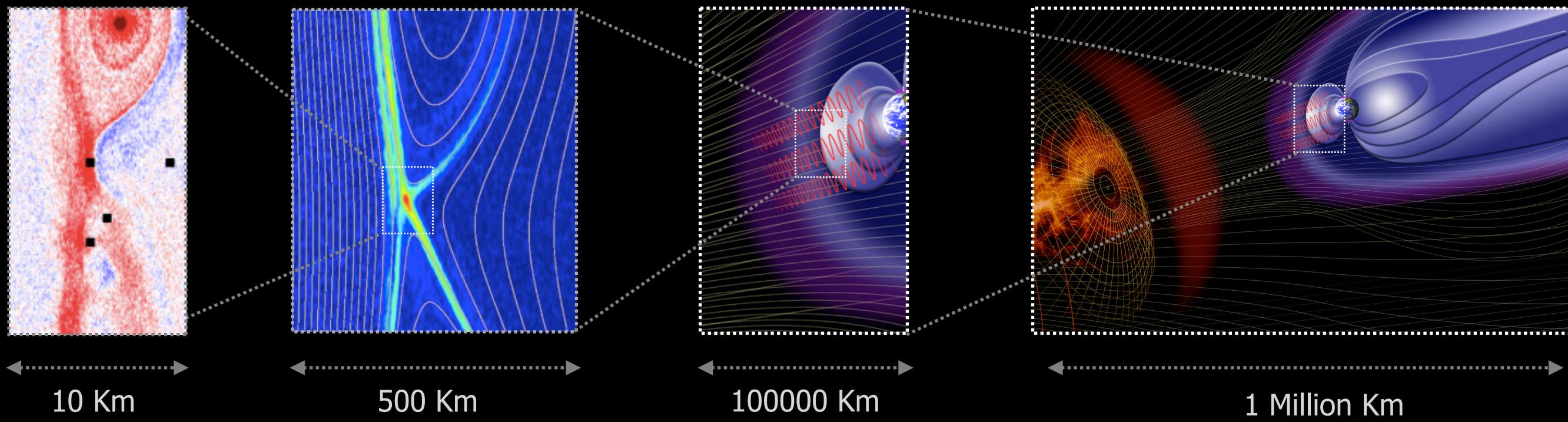
Andrea Ciardi

Roch Smets

Alexis Jeandet



THE MULTISCALE CHALLENGE

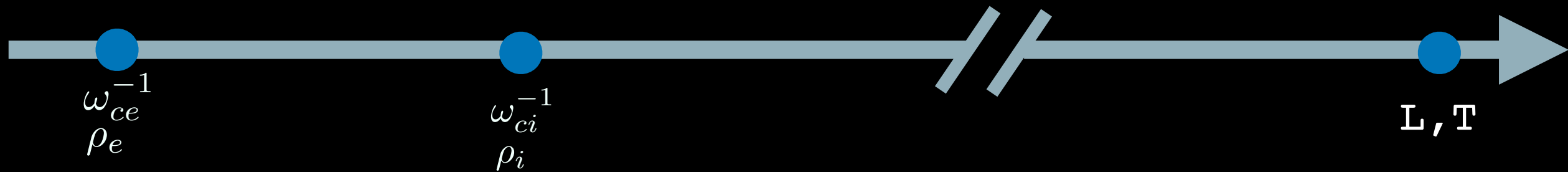


Fully Kinetic

fully kinetic physics, but small domains, short durations

Fluid « MHD »

Global scale system
But no kinetic physics

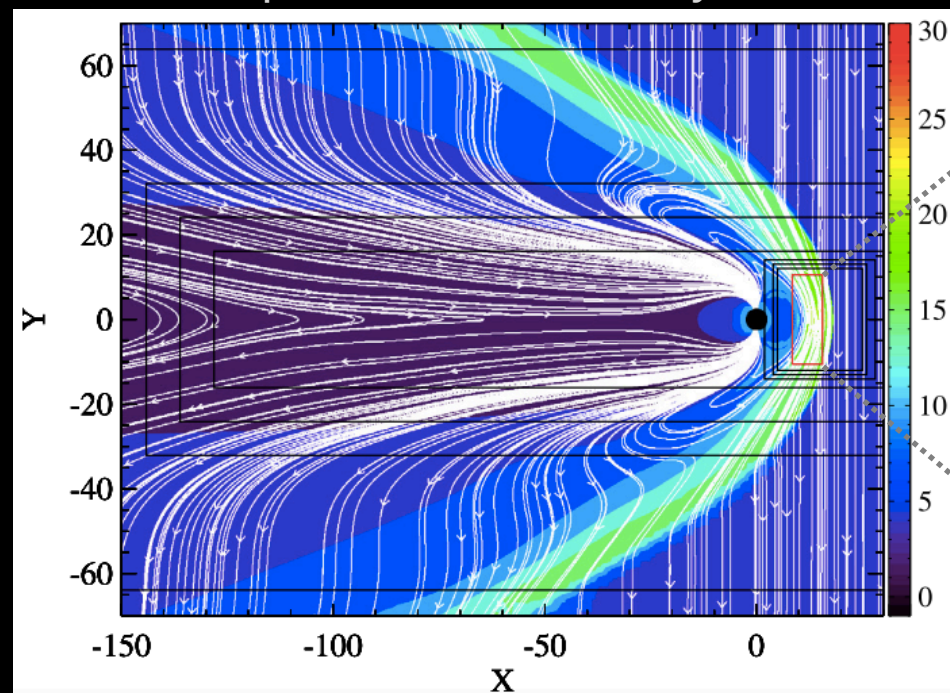


Solar corona : $\delta_e = 20cm$ $\delta_i = 10m$
(insane)

$L = 10^9 m$

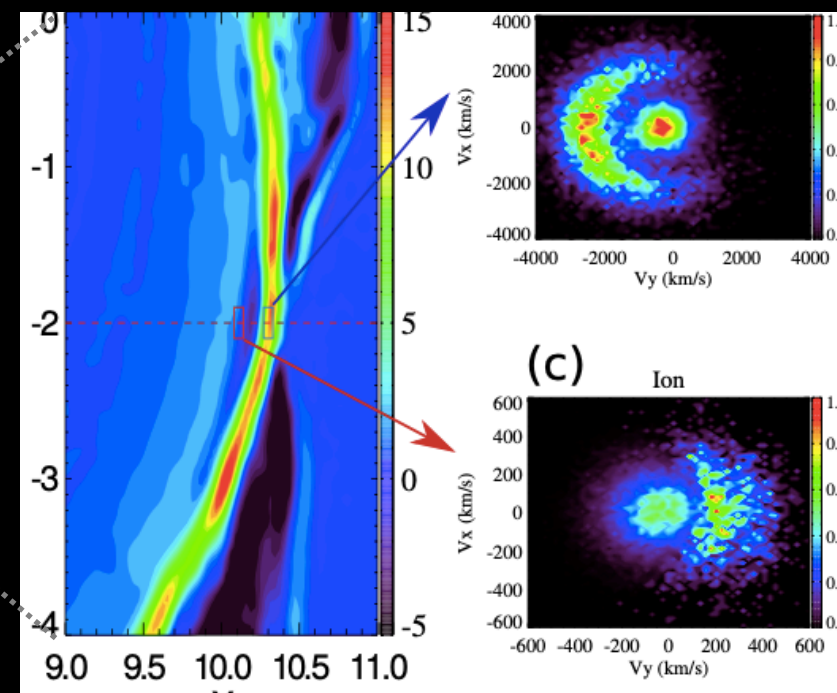
COUPLING MHD WITH FULLY KINETIC MODELS

Global MHD model coupled with implicit full PIC two-ways



[Daldorff et al. 2014]

Embedded full PIC domain with iPIC



[Yuxi Chen et al 2017]

Full-PIC



$$\omega_{ce}^{-1}$$

$$\rho_e$$



Fluid « MHD »

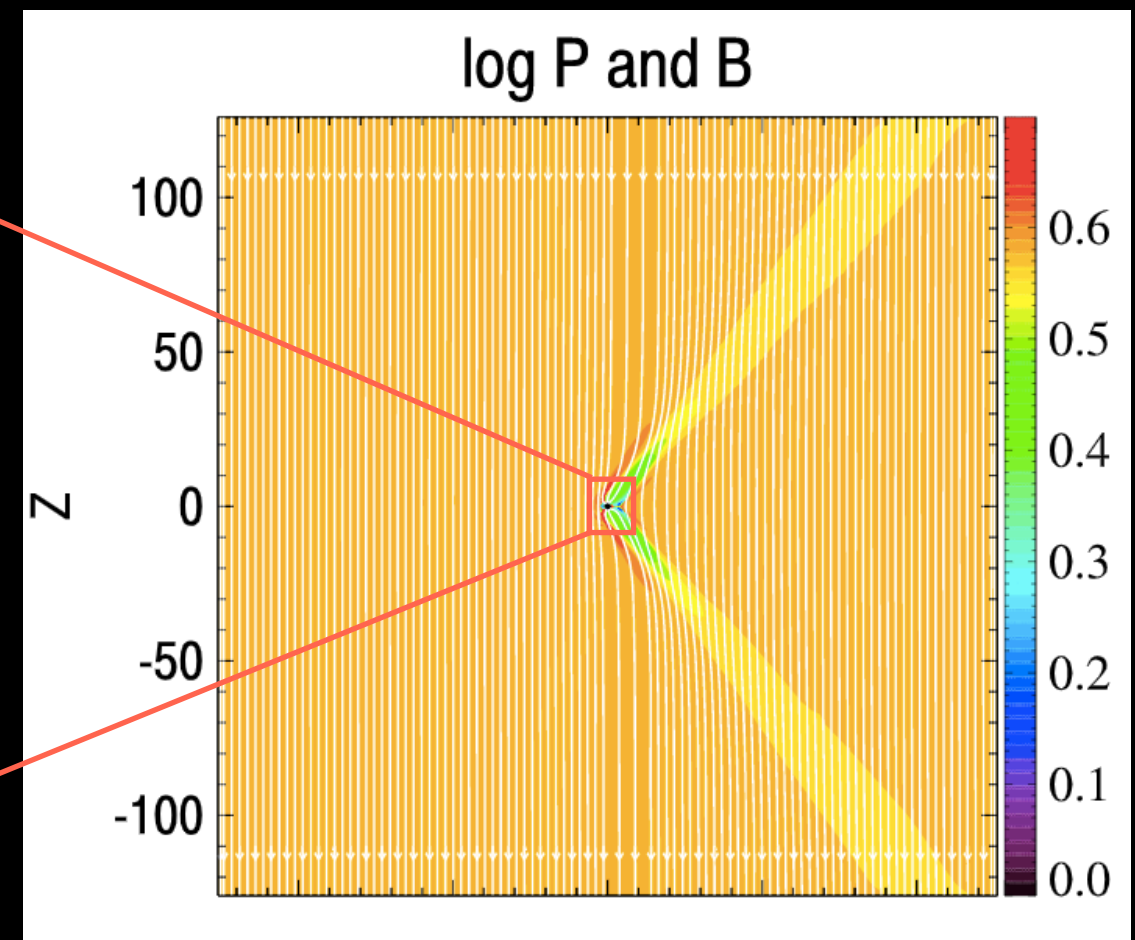
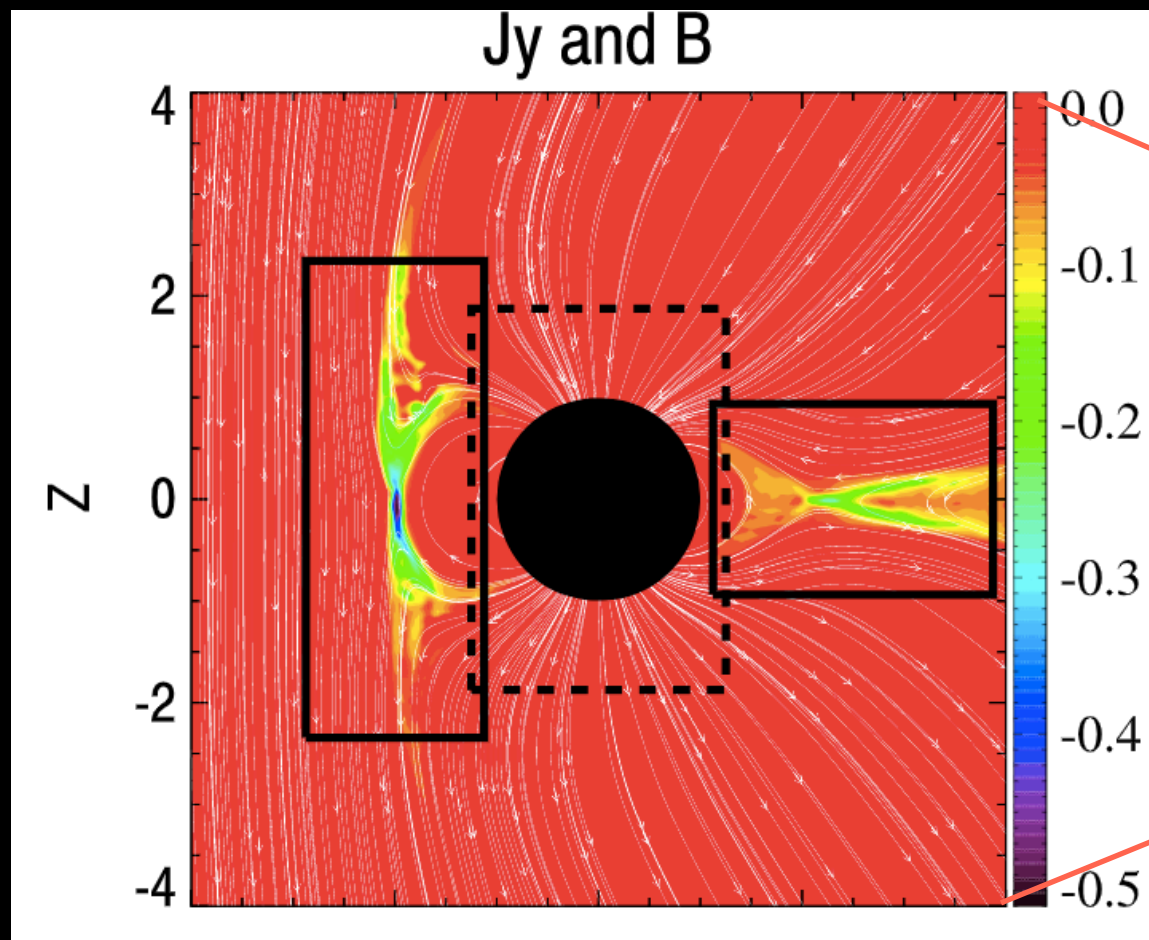


L, T

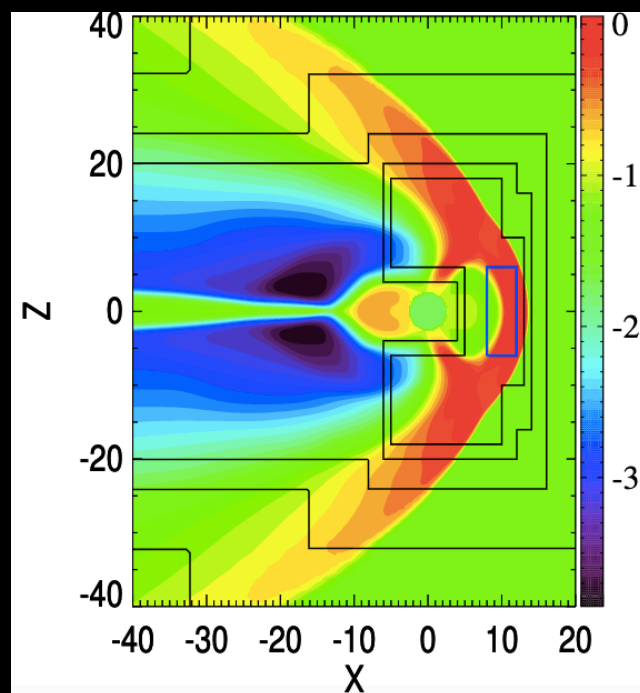
Kinetic domain position is fixed

Full-PIC is quite heavy : limited spatial domain : no global kinetic ions

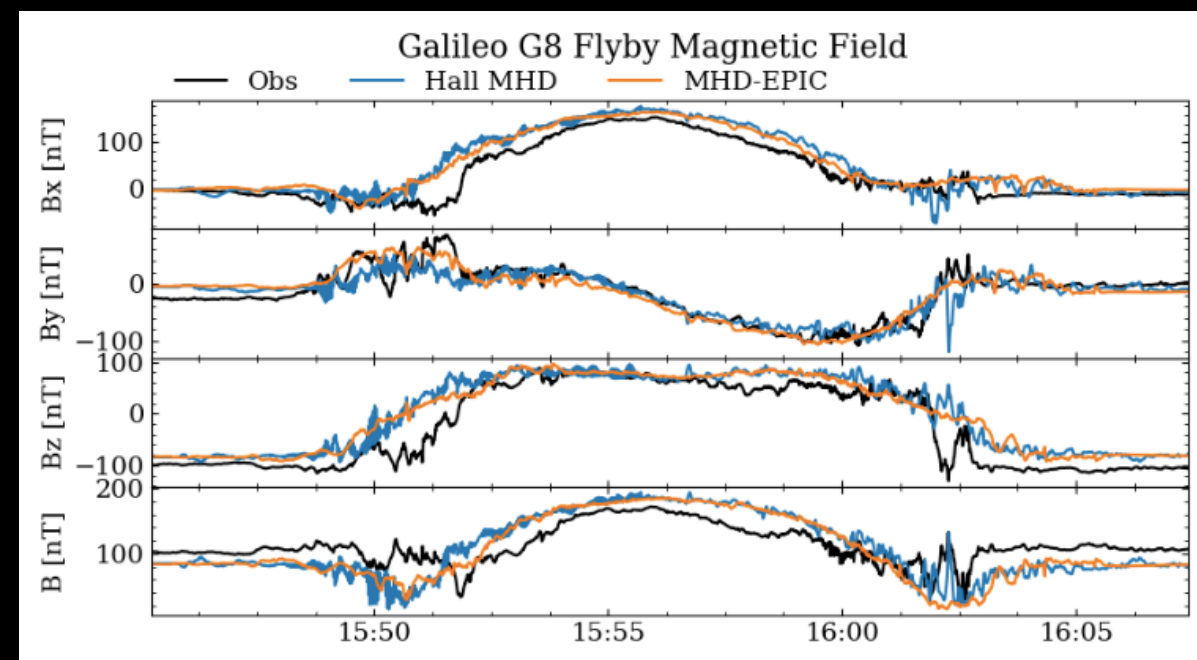
COUPLING MHD WITH FULLY KINETIC MODELS



[Toth et al 2014]

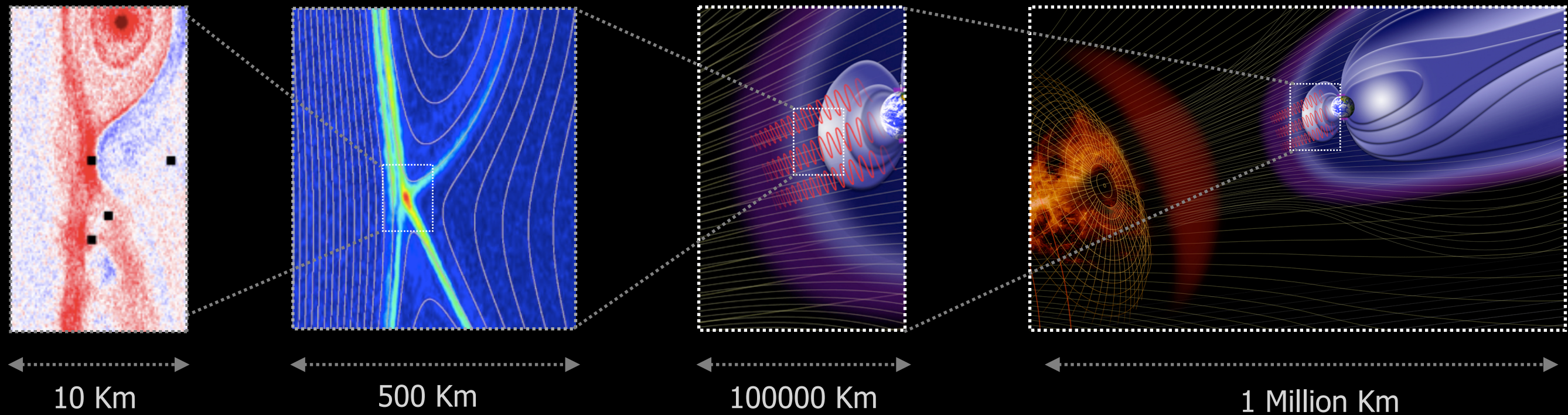


[Chen et al 2017]



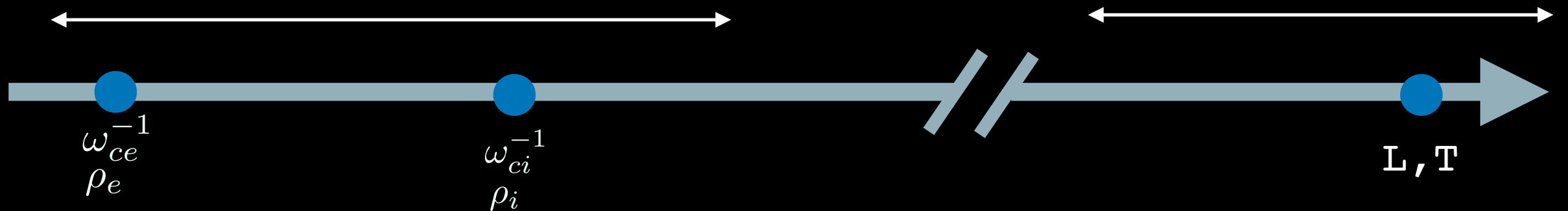
[Zhou et al 2020]

FILLING THE GAP BETWEEN SMALL AND LARGE SCALES



Fully kinetic

Fluid « MHD »



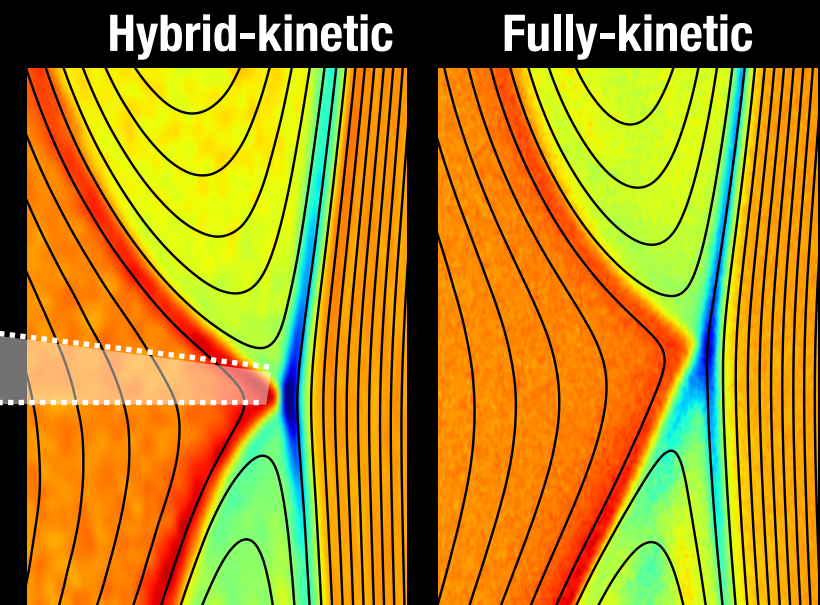
Hybrid kinetic

include ion kinetics but a fluid electron model

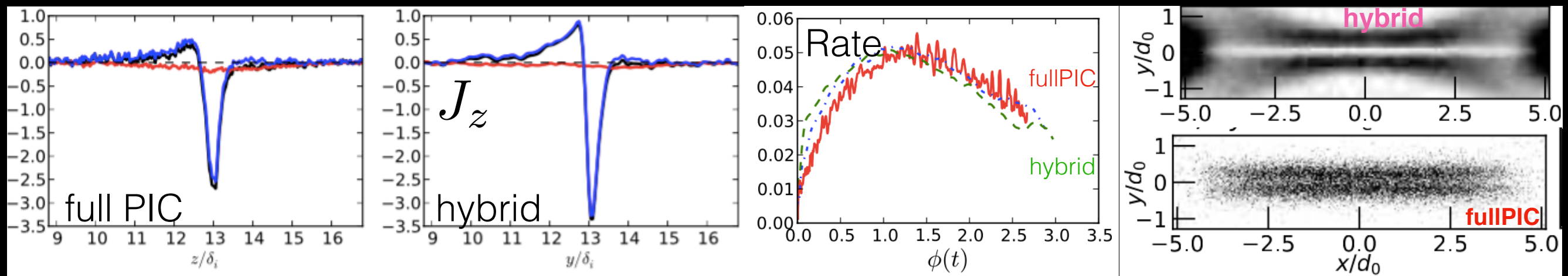
HYBRID KINETIC CAPTURE RECONNECTION DYNAMICS PRETTY WELL

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \left(-\mathbf{v}_i \times \mathbf{B} + \frac{\mathbf{j} \times \mathbf{B}}{ne} - \frac{\nabla P_e}{ne} - \nu \nabla^2 \mathbf{j} \right)$$

$\overbrace{\hspace{10em}}^{\sim 100 \text{ km}}$
 $\overbrace{\hspace{5em}}^{\sim 10 \text{ km}}$



[Aunai et al. 2013]



[Sladkov et al. 21]

Hall currents, ion acceleration/heating, collisionless mixing

Reconnection rate

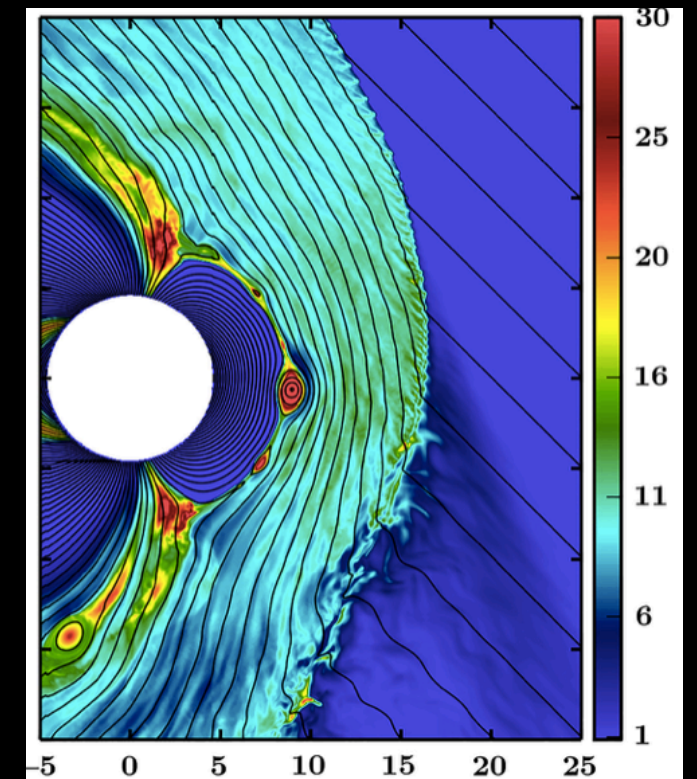
Misses X-line electron scale mechanisms, fully kinetic instabilities

Depends on the electron closure equation

$$\begin{aligned} \partial_t \mathbf{P} = & -\mathbf{V}_e \cdot \nabla \mathbf{P} - \mathbf{P} \nabla \cdot \mathbf{V}_e - \mathbf{P} \cdot \nabla \mathbf{V}_e - (\mathbf{P} \cdot \nabla \mathbf{V}_e)^T \\ & - \frac{e}{m} [\mathbf{P} \times \mathbf{B} + (\mathbf{P} \times \mathbf{B})^T] - \frac{1}{\tau} [\mathbf{P} - \frac{1}{3} \text{Tr}(\mathbf{P}) \mathbf{1}] \end{aligned}$$

GLOBAL VLASOV (EULERIAN 6D) HYBRID KINETIC MODEL

2D vlsiator run

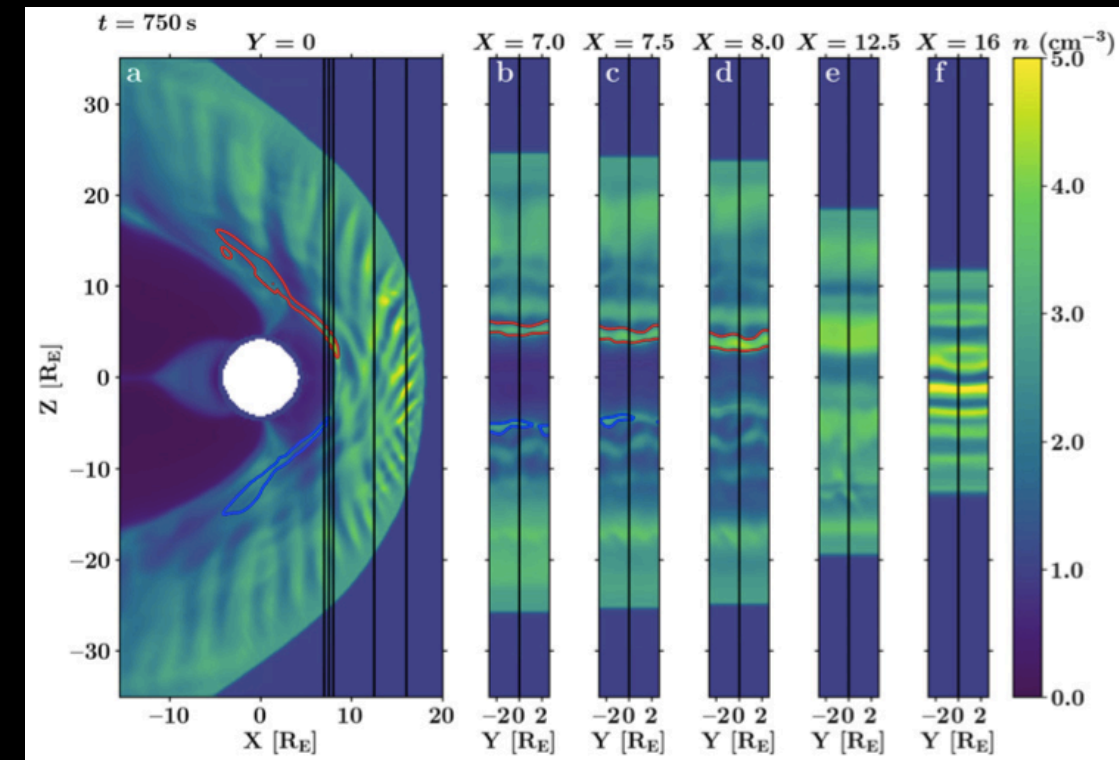
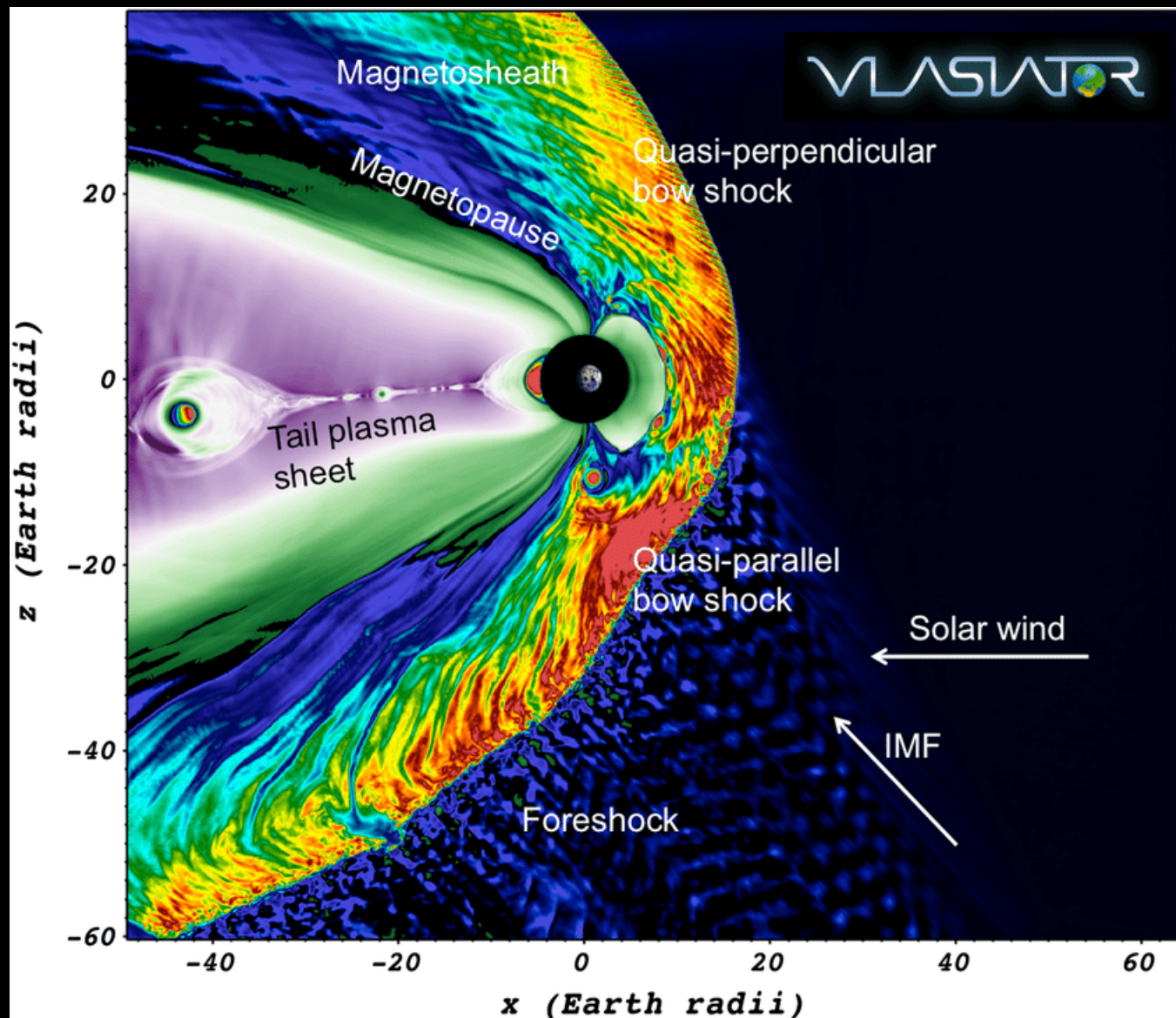


[S. Hoilijoki et al. 2017]

Started in ~2010

Advantage : noise free compared to PIC

Drawback : complex and computationally heavy : 2021 first spatial 3D

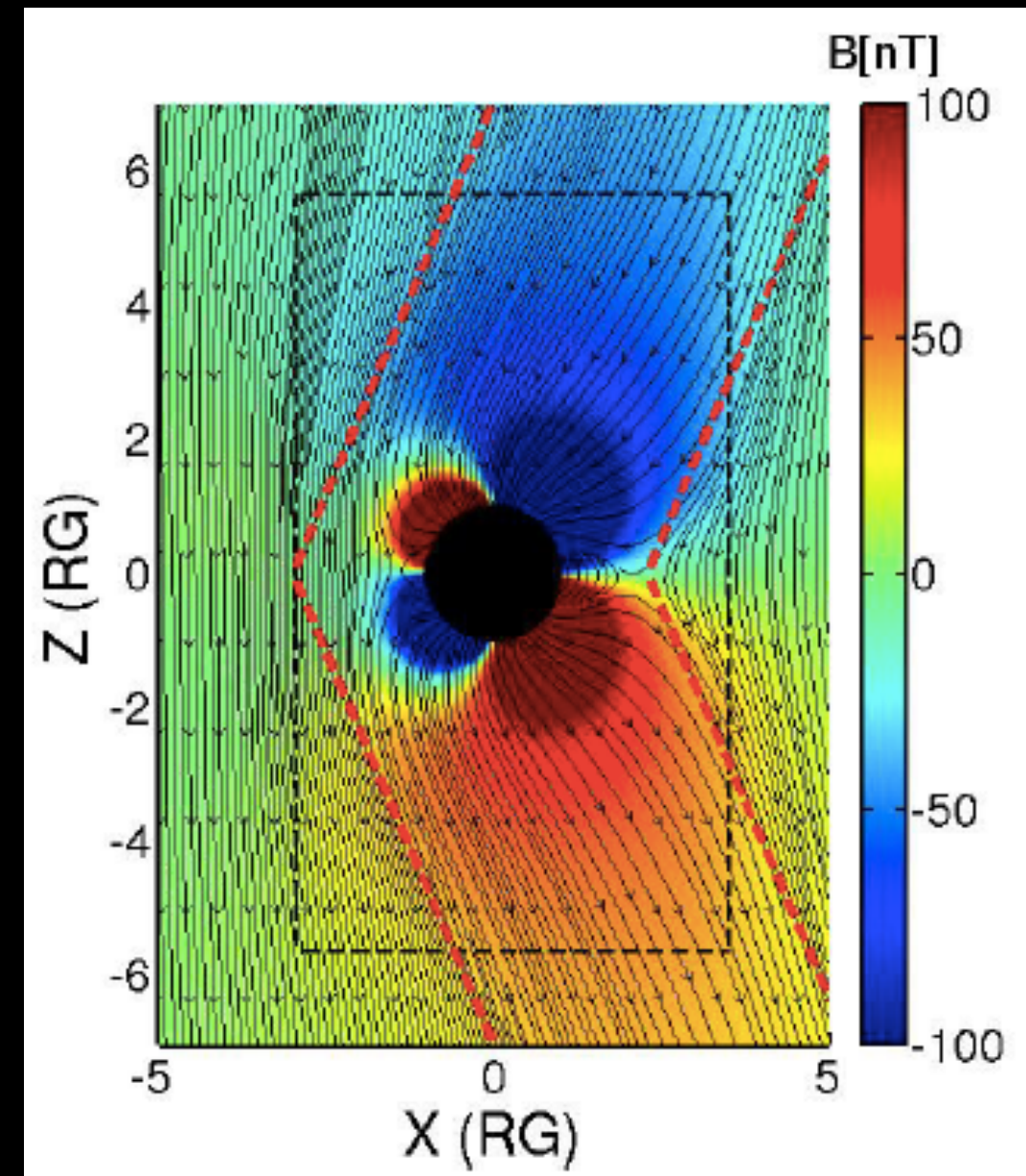
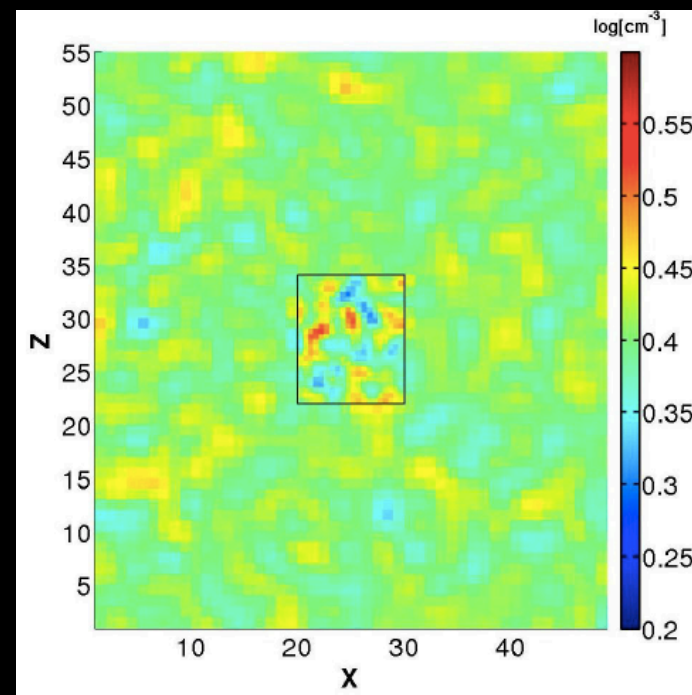
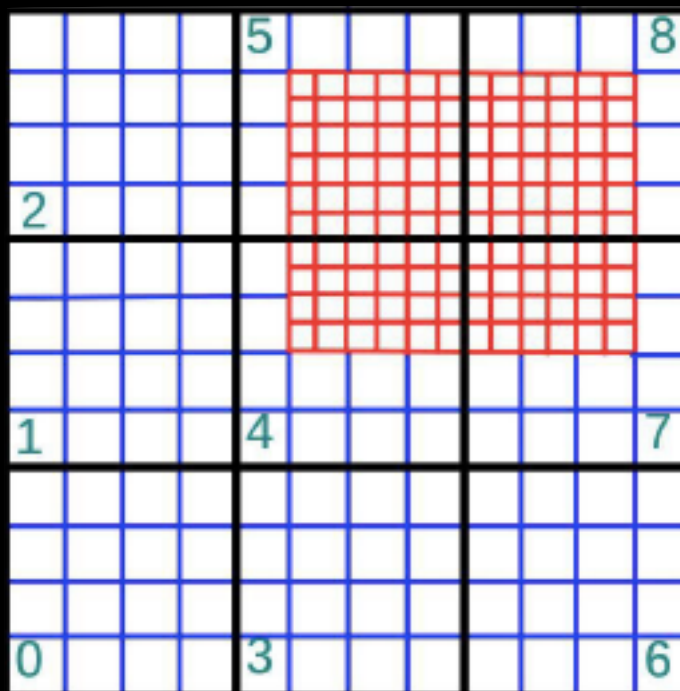


[Kempf et al. 2020]

BEYOND « CLASSICAL » HYBRID PIC CODES

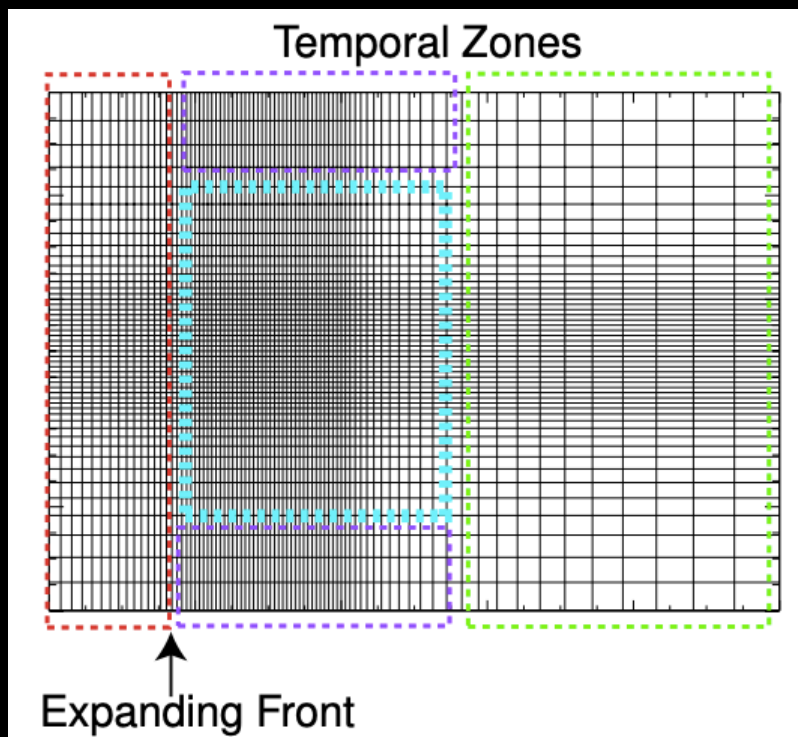
A first step : using a refined grid around a region of interest

[Leclercq et al. 2016]

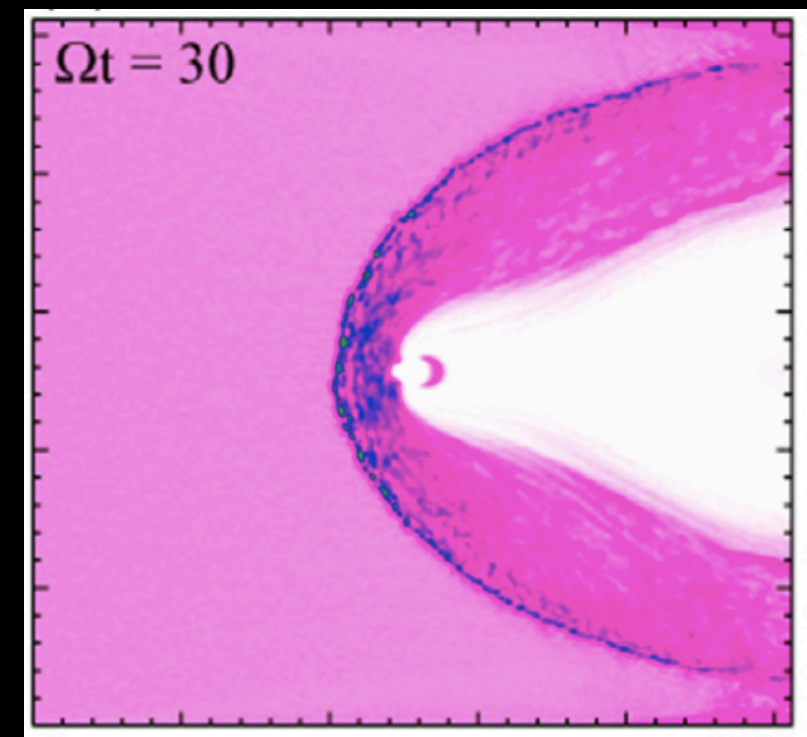
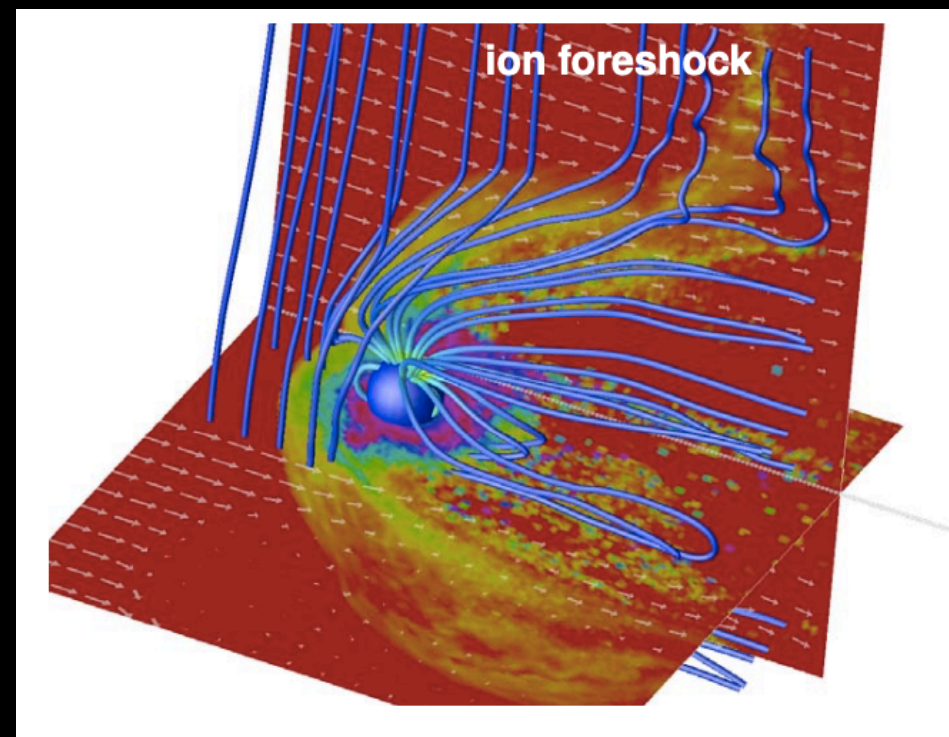


- Better than uniform grid
- *Home made* gridding
- Not adaptive

GLOBAL HYBRID PARTICLE IN CELL



[Karimabadi et al. 2006]



[Omelchenko et al. 2012]

Time zones and stretched grids can address this issue but not so flexible

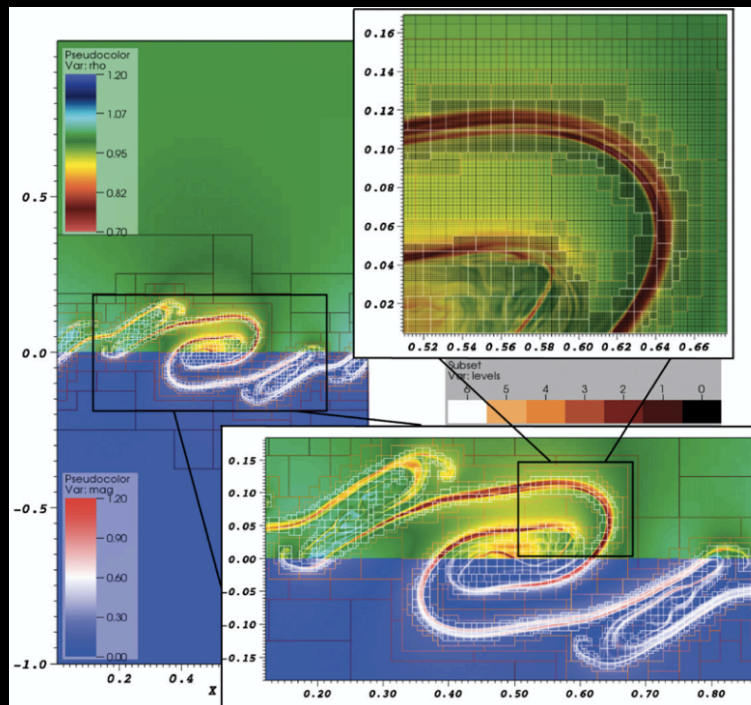
Discrete Event Simulations (DES)
Each node/particle evolved when needed

ADAPTIVE MESH REFINEMENT (AMR) (IN FLUID CODES...)

« finer mesh is more accurate, but more costly »

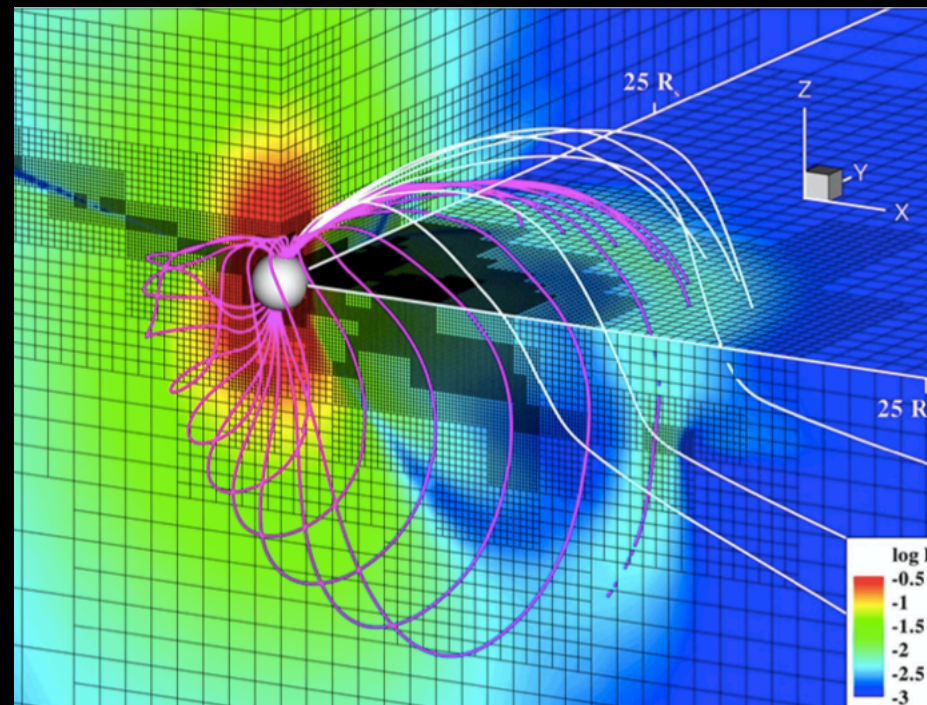
Refine the mesh where a certain set of criteria is met to provide a better accuracy of the solution in critical regions while keeping a coarser resolution in regions of less interest

PLUTO



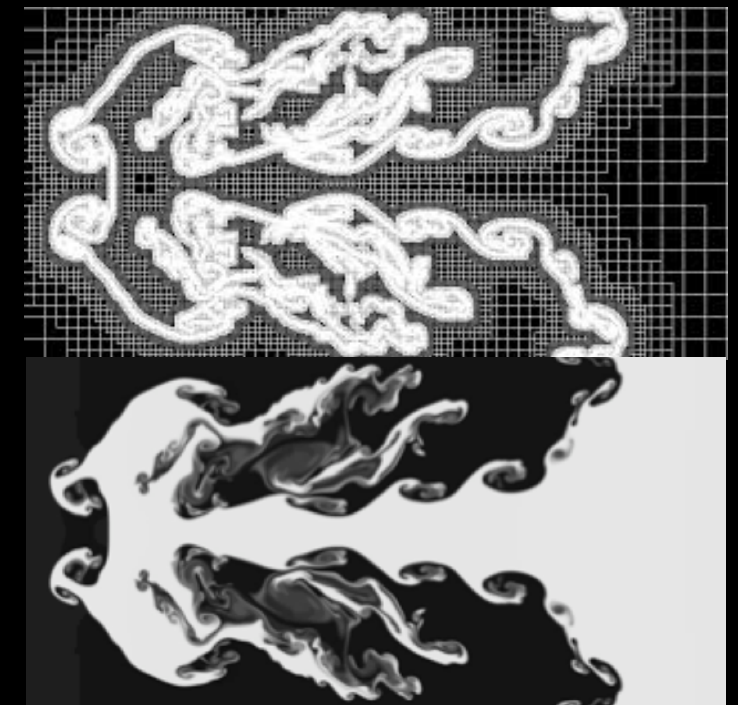
[Mignone et al. 2012]

BATSRUS



[Gombosi et al. 2003]

RAMSES

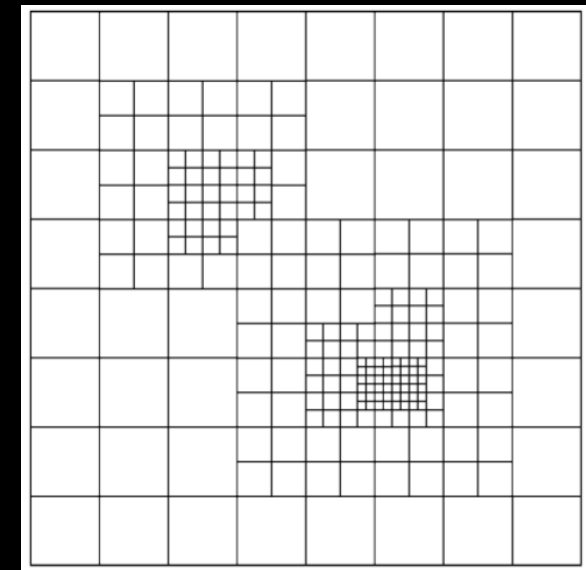
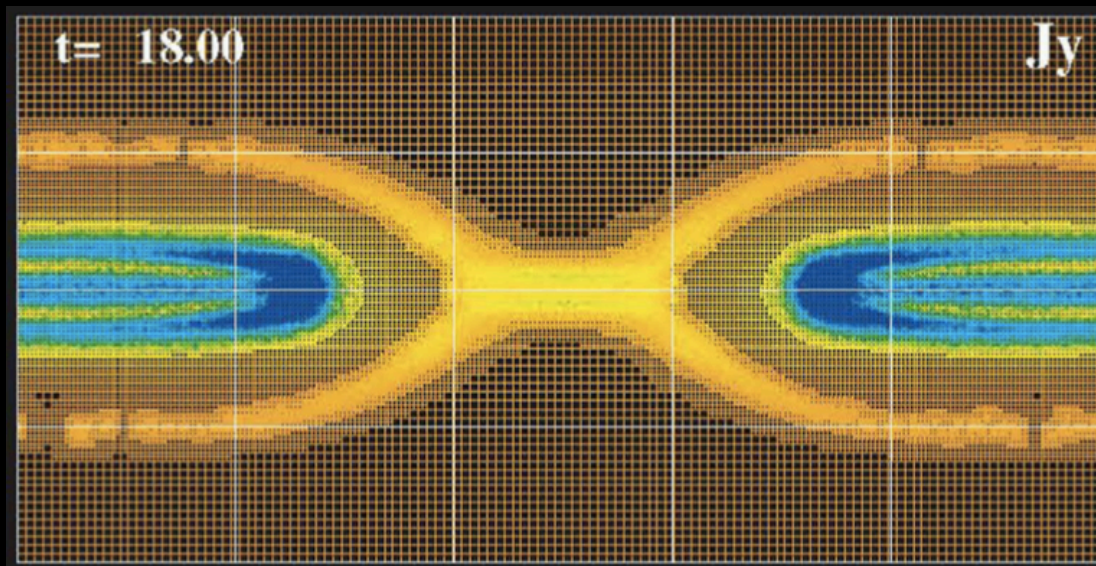


[Teyssier et al. 200..]

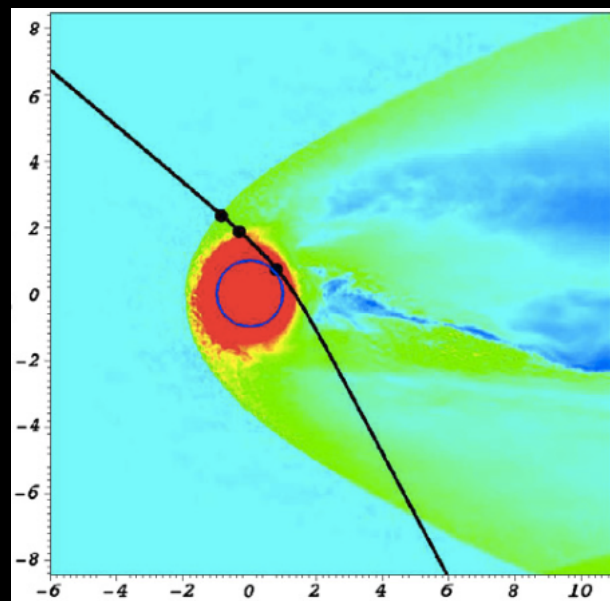
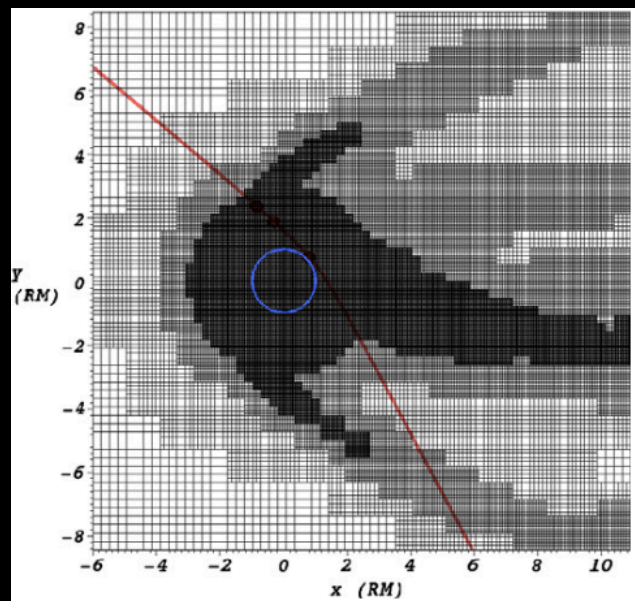
AMR is now a mainstream technology in MHD codes

AMR IS RARE IN PIC CODES

Explicit full PIC code with tree (cell)-based AMR still needs mesh Debye-scale



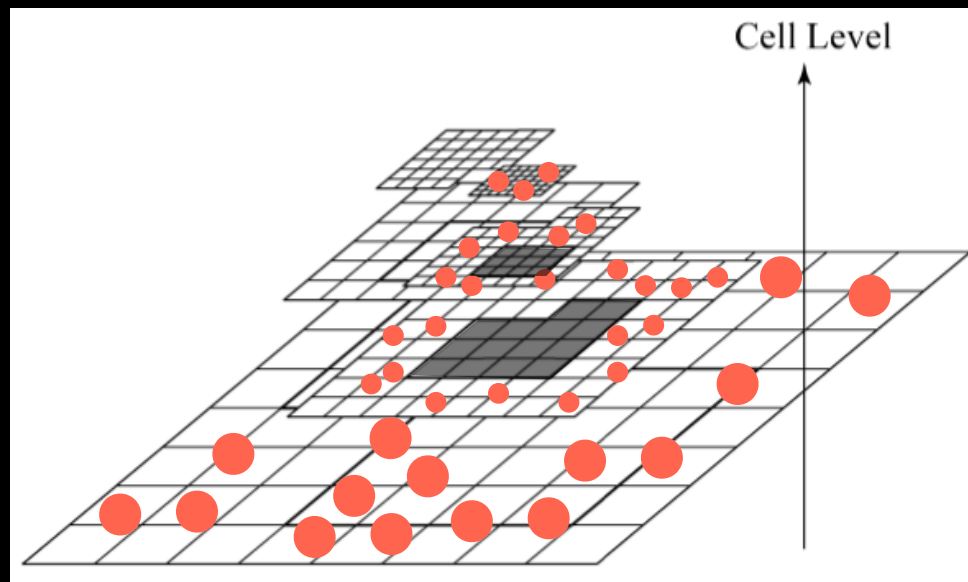
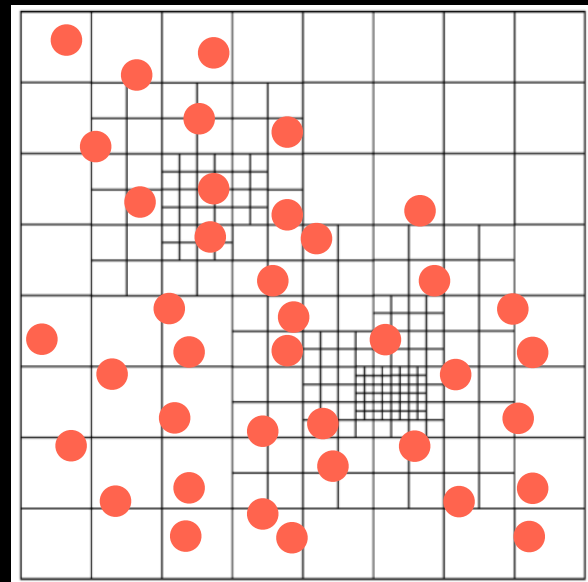
[Fujimoto et al. 2008]



[Muller et al. 2011]

Hybrid PIC with hybrid block AMR, since 2012 only used with fixed grids

AMR ONLY FOR FIELDS



Field equations are solved on each grid level

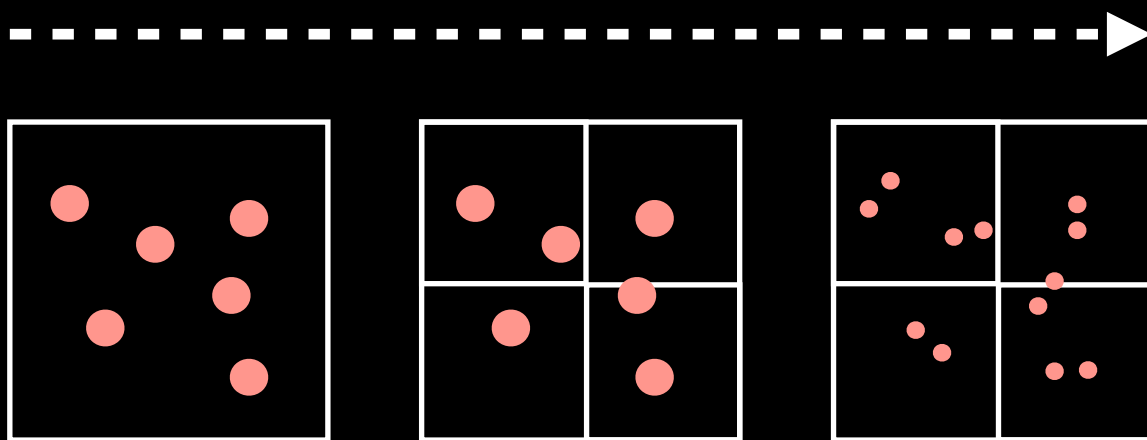
However macroparticles only interact with the finest mesh at their location

[adapted from Fujimoto et al. 2008]

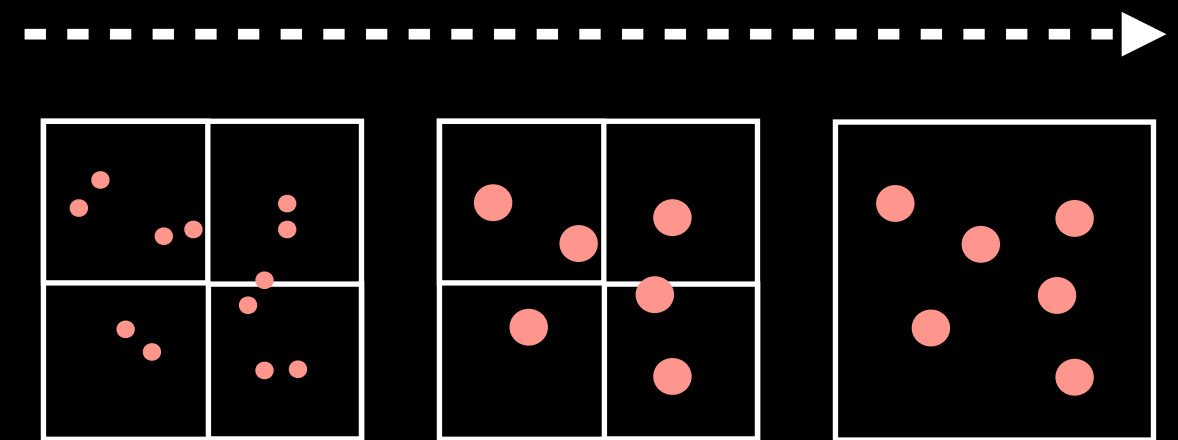
refine

Splitting

Coarsening and merging



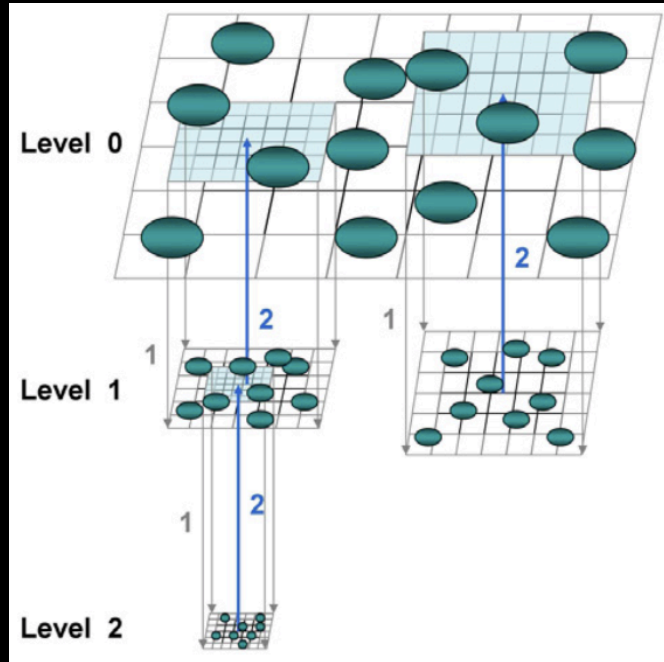
Split macroparticles for coarse to fine flux to keep noise constant



Merging does not conserve the distribution function...

MULTI-DOMAIN-MULTI-LEVEL (MLMD) SOLVES AMR PIC ISSUES

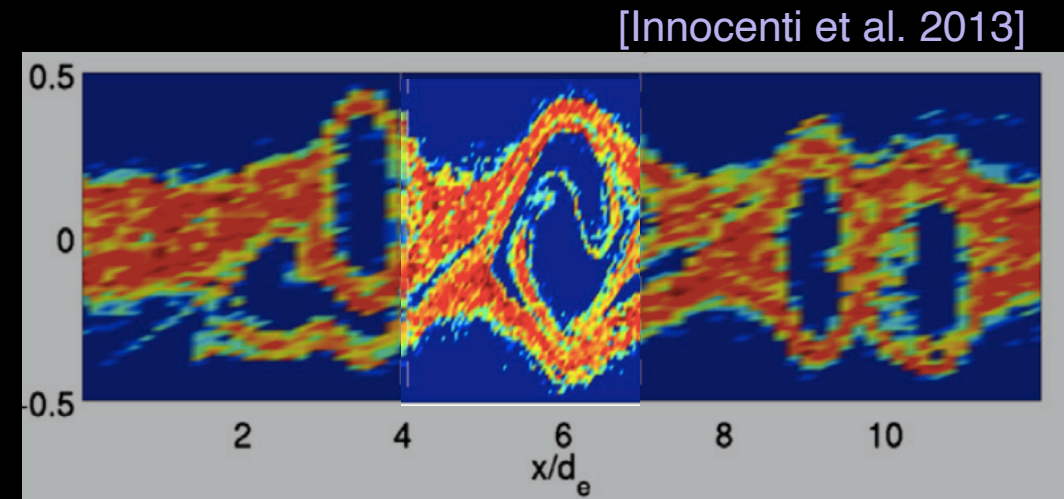
Solve all equations (like fluid) on all levels : fields AND particles.
No need to merge macroparticles!



[Innocenti et al. 2013]

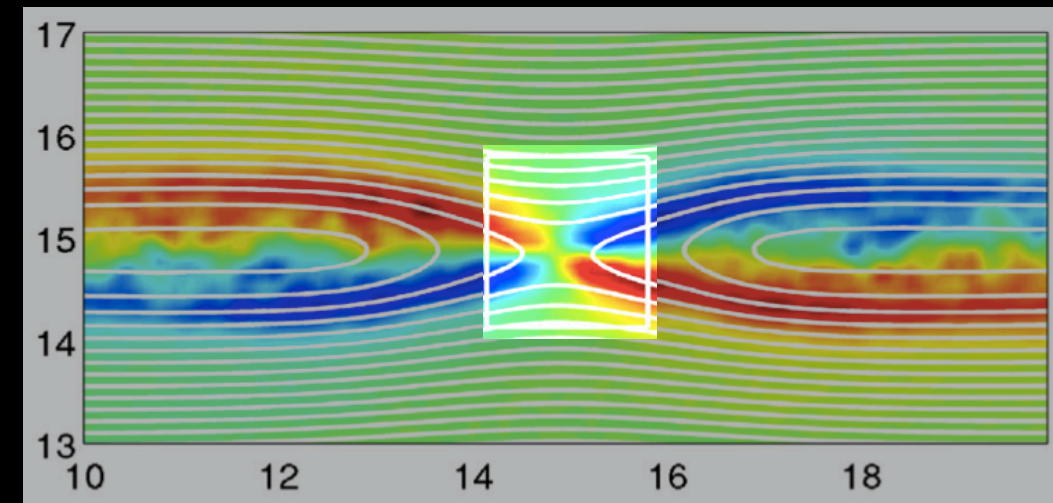
Levels are coupled through boundaries and synchronization but otherwise rather independent : opens the way to multi-formalisms.

Concept in 1D with implicit PIC



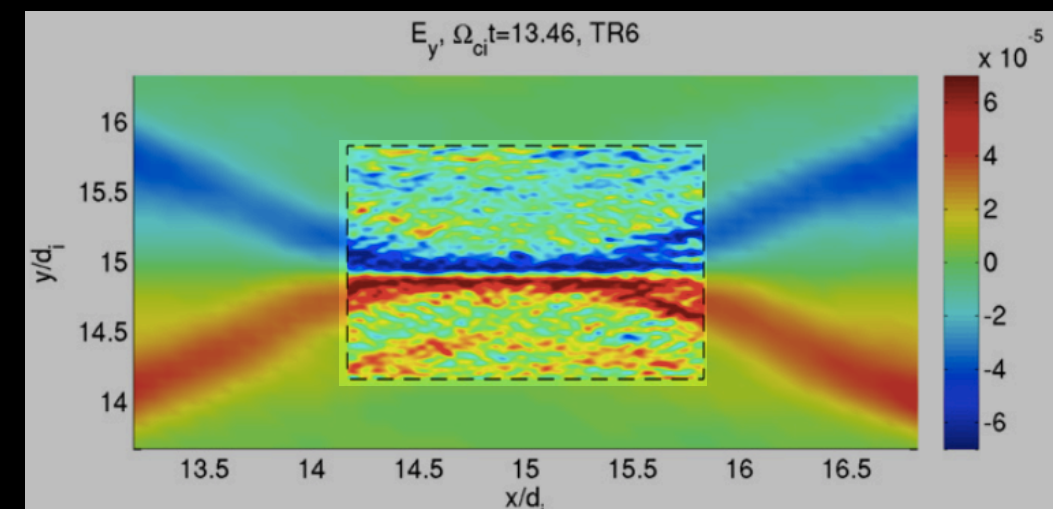
[Innocenti et al. 2013]

Generalized to 2D



[Beck et al. 2014]

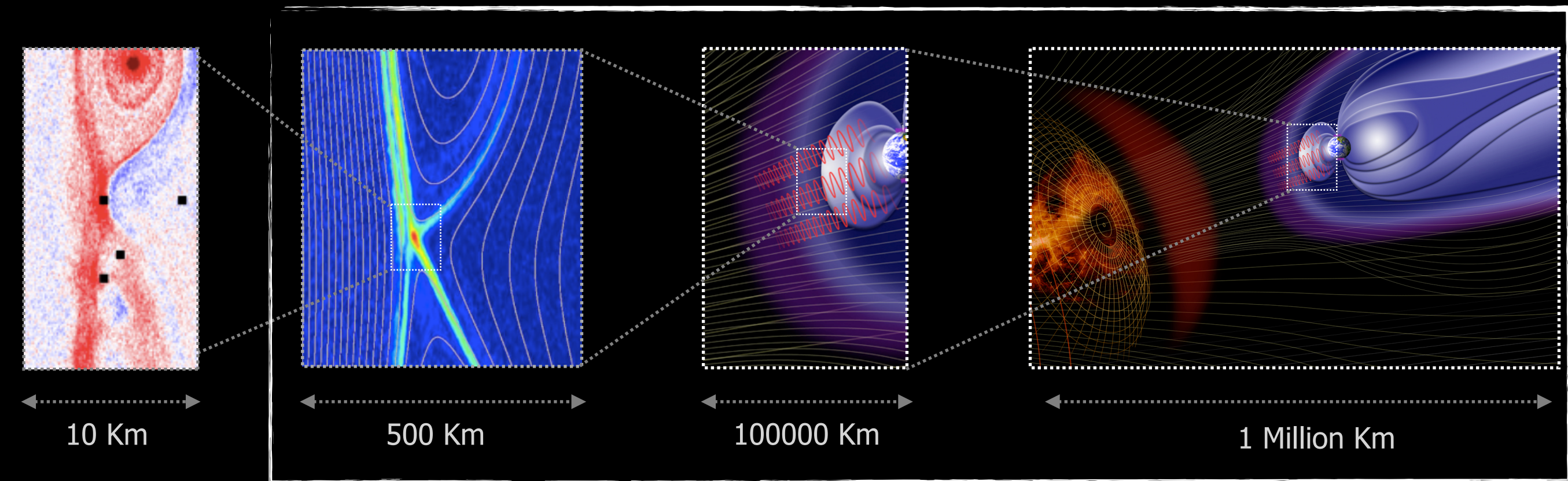
Time subcycling



[Innocenti et al. 2015]

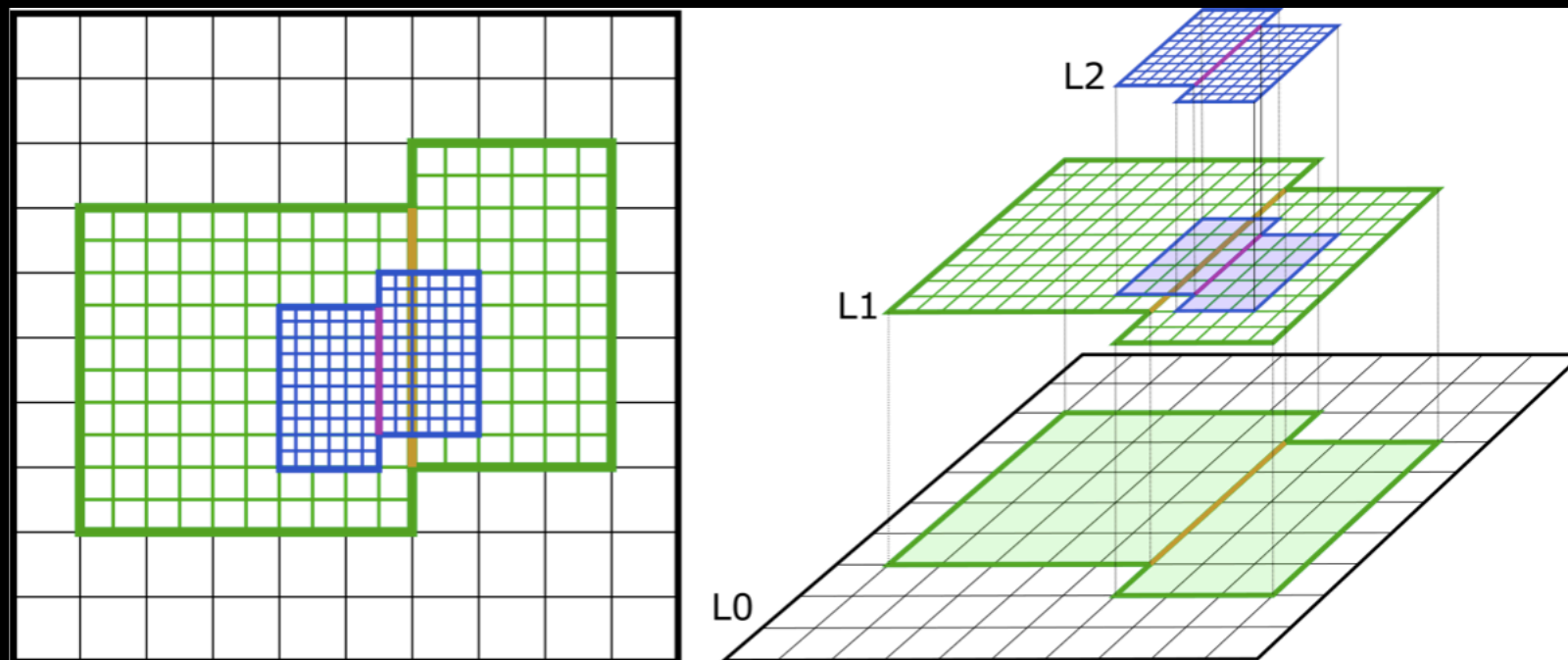
Unfortunately : home-made, limited to 1 (fixed) refined level (but with refined ratio > 2)

PHARE: AMR HYBRID-PIC CODE



PHARE

Structured
Adaptive Mesh
Refinement



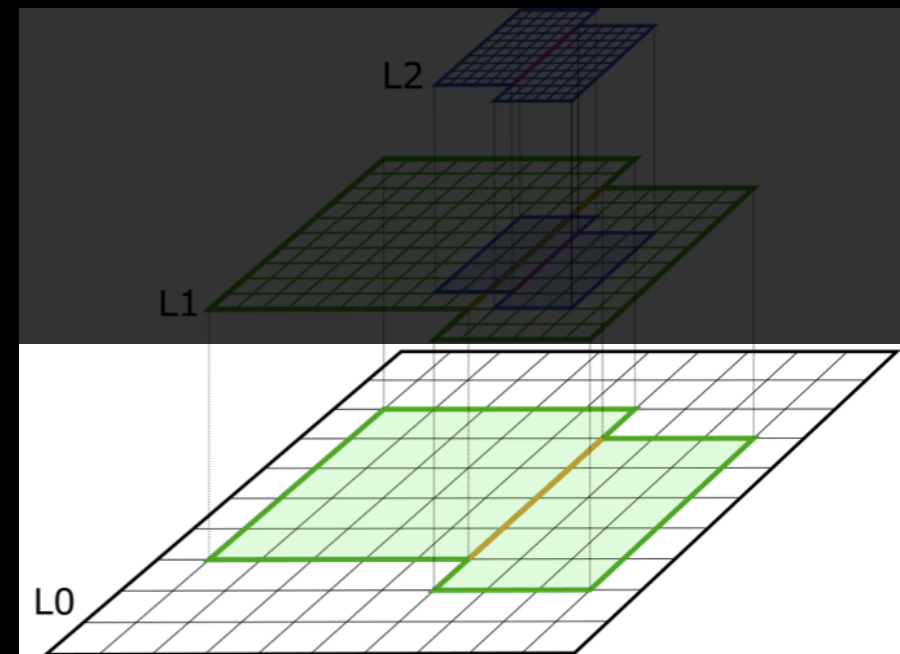
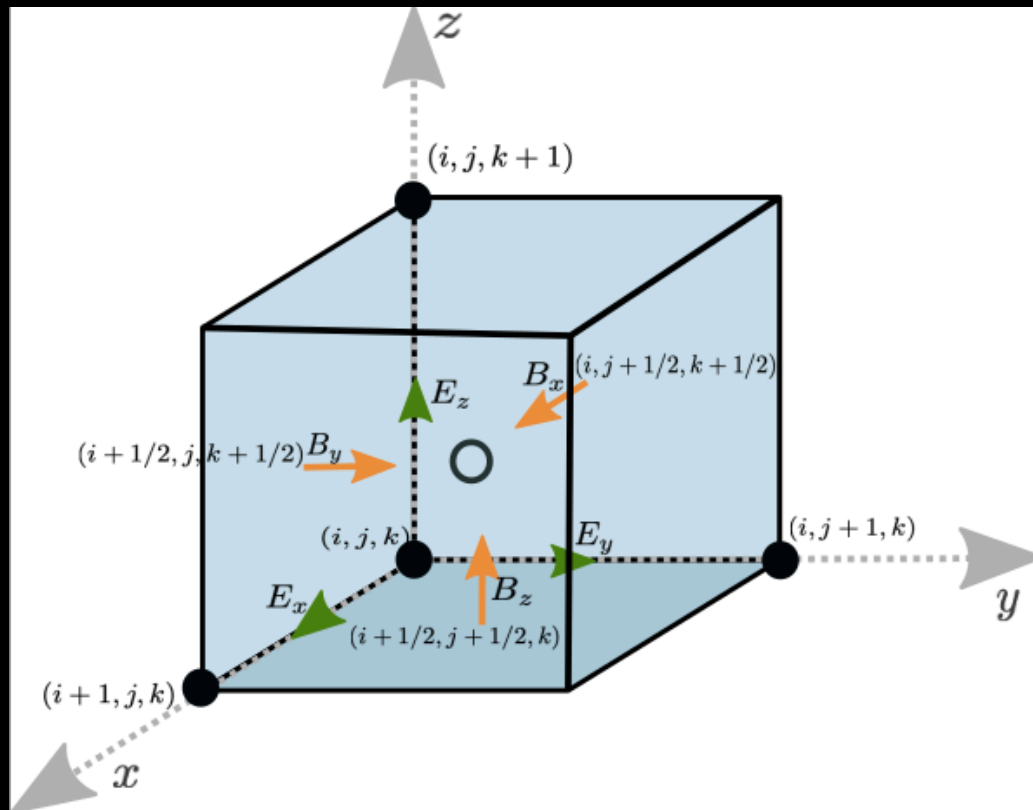
AMR
Hybrid-PIC

Hybrid or
MHD?

SPATIAL DISCRETIZATION

Equations are discretized following the Yee layout [Yee 1966]

This conserves $\nabla \cdot \mathbf{B} = 0$ intrinsically



t

Prediction

$$\mathbf{B}_{p1}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^n$$

$$\mathbf{E}_{p1}^{n+1} = -\mathbf{u}^n \times \mathbf{B}_{p1}^{n+1} + \frac{\nabla \times \mathbf{B}_{p1}^{n+1}}{N^n} - \frac{\nabla \cdot \mathbf{P}_e}{N^n} + \eta \nabla \times \mathbf{B}_{p1}^{n+1} - \nu \nabla^2 \nabla \times \mathbf{B}_{p1}^{n+1}$$

$$(\mathbf{E}, \mathbf{B})^{n+1/2} = \langle (\mathbf{E}, \mathbf{B}) \rangle_n^{n+1}$$

$$\mathbf{r}_{p1}^{n+1/2} = \mathbf{r}^n + \Delta t / 2 \mathbf{v}^n$$

$$\mathbf{E}, \mathbf{B}(\mathbf{r}_{p1}^{n+1/2}) = \sum_{ijk} (\mathbf{E}, \mathbf{B}_{ijk}) W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1/2}|)$$

$$m_i \frac{d\mathbf{v}_{p1}^{n+1}}{dt} = e \left(\mathbf{v}^n \times + \mathbf{B}^{n+1/2} + \mathbf{E}^{n+1/2} \right)$$

$$N^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|) \quad u^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|)$$

Prediction

$$\mathbf{B}_{p2}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^{n+1/2}$$

$$\mathbf{E}_{p2}^{n+1} = -\mathbf{u}^{n+1} \times \mathbf{B}_{p2}^{n+1} + \frac{\nabla \times \mathbf{B}_{p2}^{n+1}}{N^{n+1}} - \frac{\nabla \cdot \mathbf{P}_e}{N^{n+1}} + \eta \nabla \times \mathbf{B}_{p2}^{n+1} - \nu \nabla^2 \nabla \times \mathbf{B}_{p2}^{n+1}$$

$$\mathbf{r}_{p2}^{n+1/2} = \mathbf{r}^n + \Delta t / 2 \mathbf{v}^n$$

$$(\mathbf{E}, \mathbf{B})^{n+1/2} = \langle (\mathbf{E}, \mathbf{B}) \rangle_n^{n+1}$$

$$m_i \frac{d\mathbf{v}_{p2}^{n+1}}{dt} = e \left(\mathbf{v}^n \times + \mathbf{B}^{n+1/2} + \mathbf{E}^{n+1/2} \right)$$

$$N^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|) \quad u^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|)$$

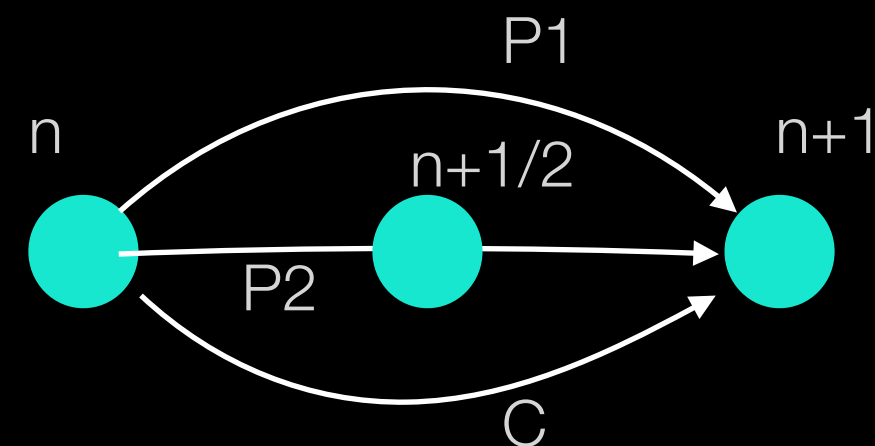
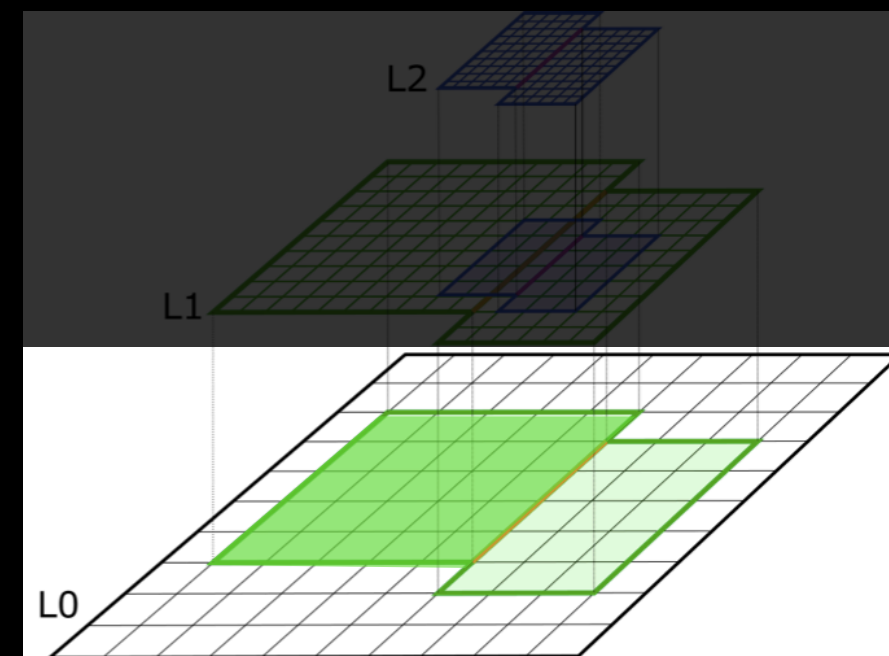
Correction

$$\mathbf{B}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^{n+1/2}$$

$$\mathbf{E}^{n+1} = -\mathbf{u}^{n+1} \times \mathbf{B}^{n+1} + \frac{\nabla \times \mathbf{B}^{n+1}}{N^{n+1}} - \frac{\nabla \cdot \mathbf{P}_e}{N^{n+1}} + \eta \nabla \times \mathbf{B}^{n+1} - \nu \nabla^2 \nabla \times \mathbf{B}^{n+1}$$

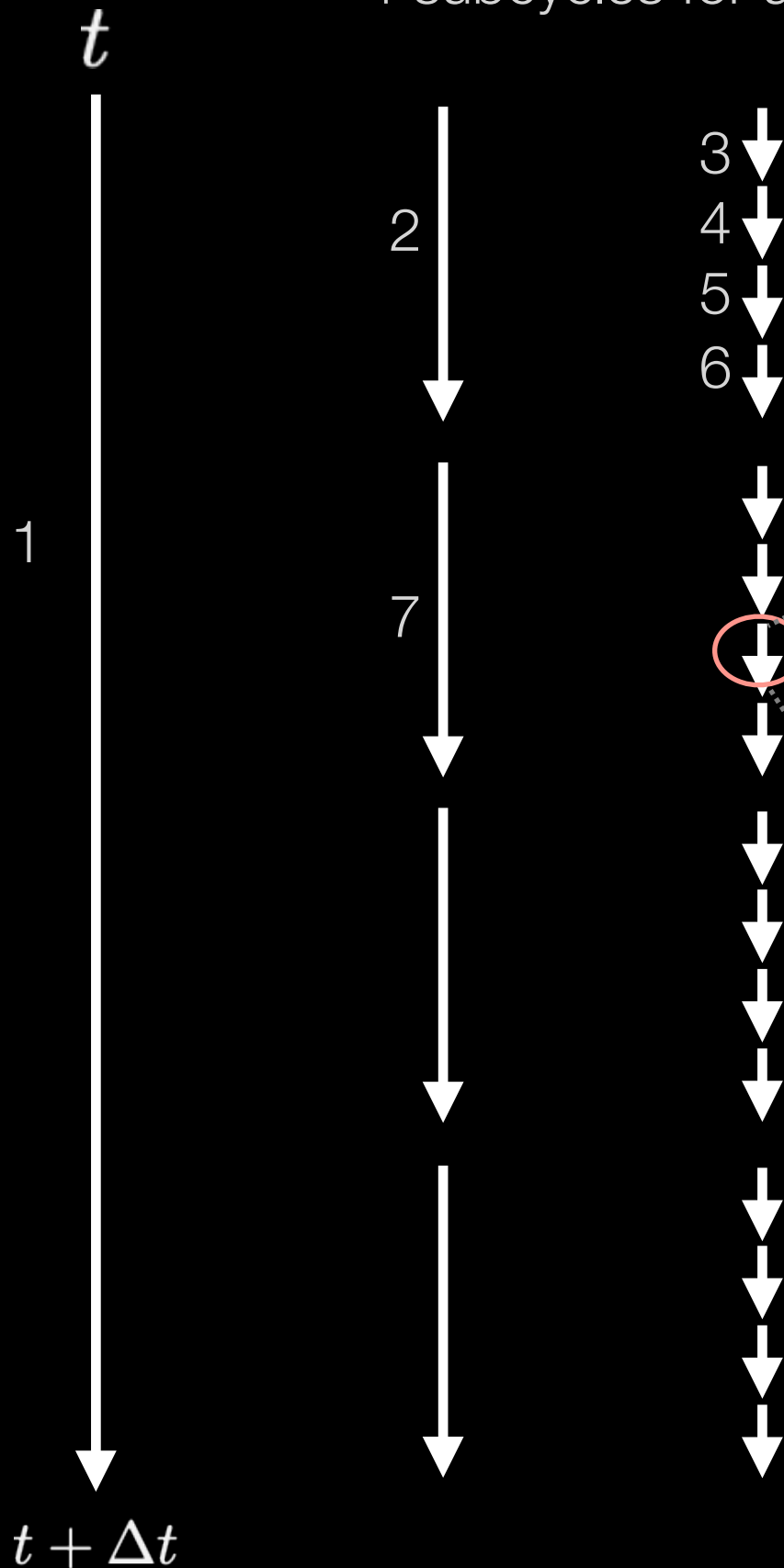
$t + \Delta t$

Iterated Crank-Nicholson
Predictor-Predictor-Corrector (PPC)
[Kunz al. 2014]



RECURSIVE TIME STEPPING

4 subcycles for a x2 refinement ratio : whistler wave CFL



PPC step

Prediction

$$\mathbf{B}_{p1}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^n$$

$$\mathbf{E}_{p1}^{n+1} = -\mathbf{u}^n \times \mathbf{B}_{p1}^{n+1} + \frac{\nabla \times \mathbf{B}_{p1}^{n+1}}{N^n} - \frac{\nabla \cdot \mathbf{P}_e}{N^n} + \eta \nabla \times \mathbf{B}_{p1}^{n+1} - \nu \nabla^2 \nabla \times \mathbf{B}_{p1}^{n+1}$$

$$(\mathbf{E}, \mathbf{B})^{n+1/2} = \langle (\mathbf{E}, \mathbf{B}) \rangle_n^{n+1}$$

$$\mathbf{r}_{p1}^{n+1/2} = \mathbf{r}^n + \Delta t / 2 \mathbf{v}^n$$

$$\mathbf{E}, \mathbf{B}(\mathbf{r}_{p1}^{n+1/2}) = \sum_{ijk} (\mathbf{E}, \mathbf{B}_{ijk}) W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1/2}|)$$

$$m_i \frac{d\mathbf{v}_{p1}^{n+1}}{dt} = e (\mathbf{v}^n \times + \mathbf{B}^{n+1/2} + \mathbf{E}^{n+1/2})$$

$$N^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|) \quad u^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|)$$

Prediction

$$\mathbf{B}_{p2}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^{n+1/2}$$

$$\mathbf{E}_{p2}^{n+1} = -\mathbf{u}^{n+1} \times \mathbf{B}_{p2}^{n+1} + \frac{\nabla \times \mathbf{B}_{p2}^{n+1}}{N^{n+1}} - \frac{\nabla \cdot \mathbf{P}_e}{N^{n+1}} + \eta \nabla \times \mathbf{B}_{p2}^{n+1} - \nu \nabla^2 \nabla \times \mathbf{B}_{p2}^{n+1}$$

$$\mathbf{r}_{p2}^{n+1/2} = \mathbf{r}^n + \Delta t / 2 \mathbf{v}^n$$

$$(\mathbf{E}, \mathbf{B})^{n+1/2} = \langle (\mathbf{E}, \mathbf{B}) \rangle_n^{n+1}$$

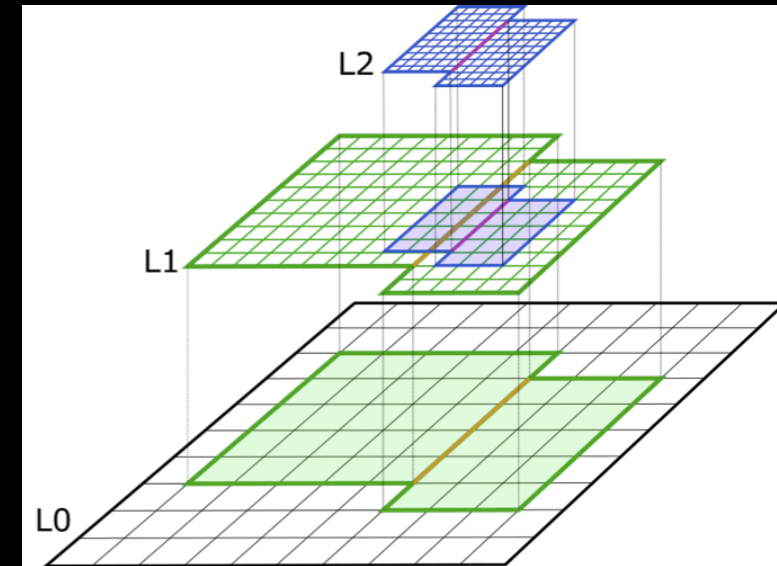
$$m_i \frac{d\mathbf{v}_{p2}^{n+1}}{dt} = e (\mathbf{v}^n \times + \mathbf{B}^{n+1/2} + \mathbf{E}^{n+1/2})$$

$$N^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|) \quad u^{n+1} = \sum_p w_p \mathbf{v}_{p1}^{n+1} W(|\mathbf{r}_{ijk} - \mathbf{r}_{p1}^{n+1}|)$$

Correction

$$\mathbf{B}^{n+1} = \mathbf{B}^n - \Delta t \nabla \times \mathbf{E}^{n+1/2}$$

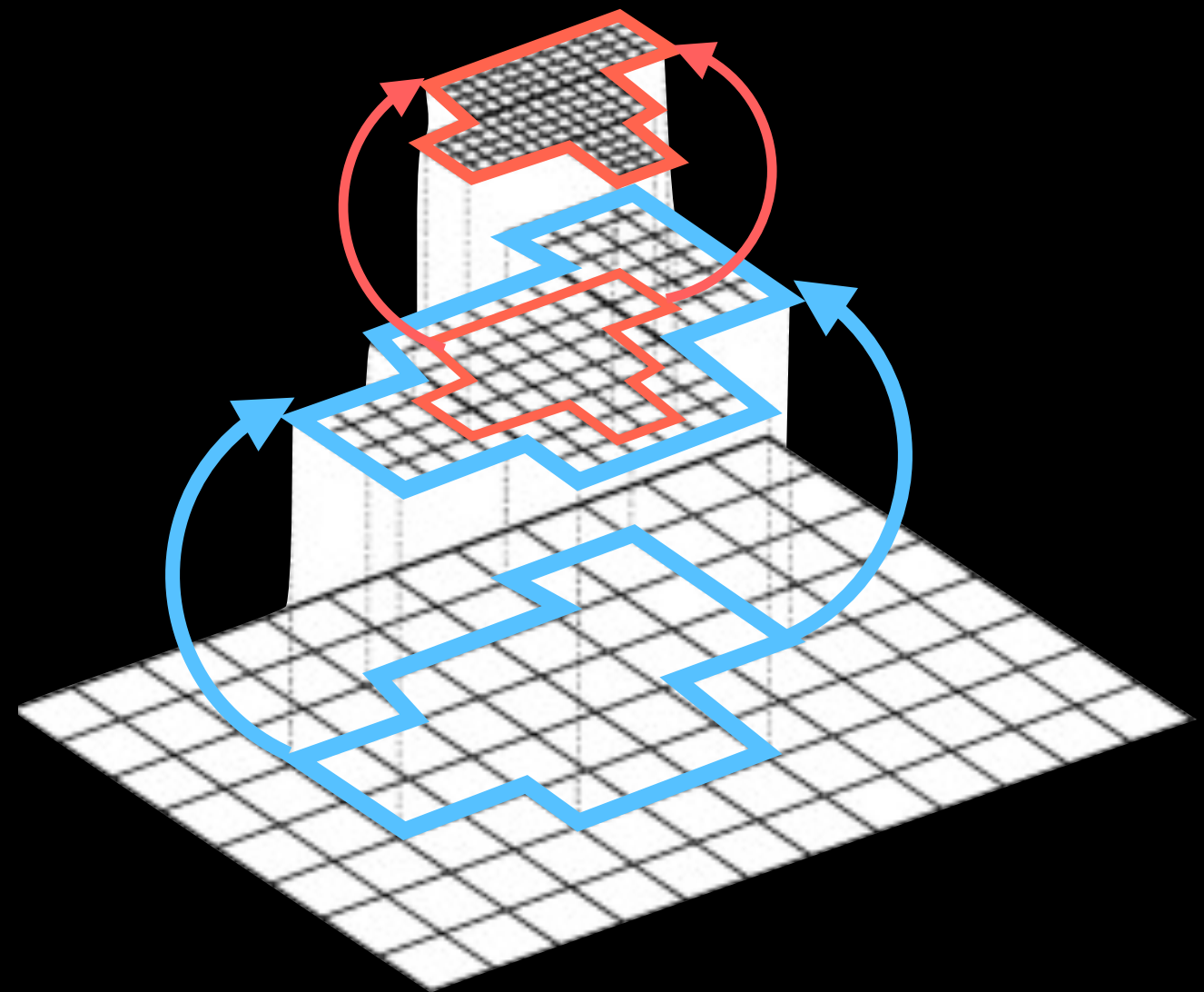
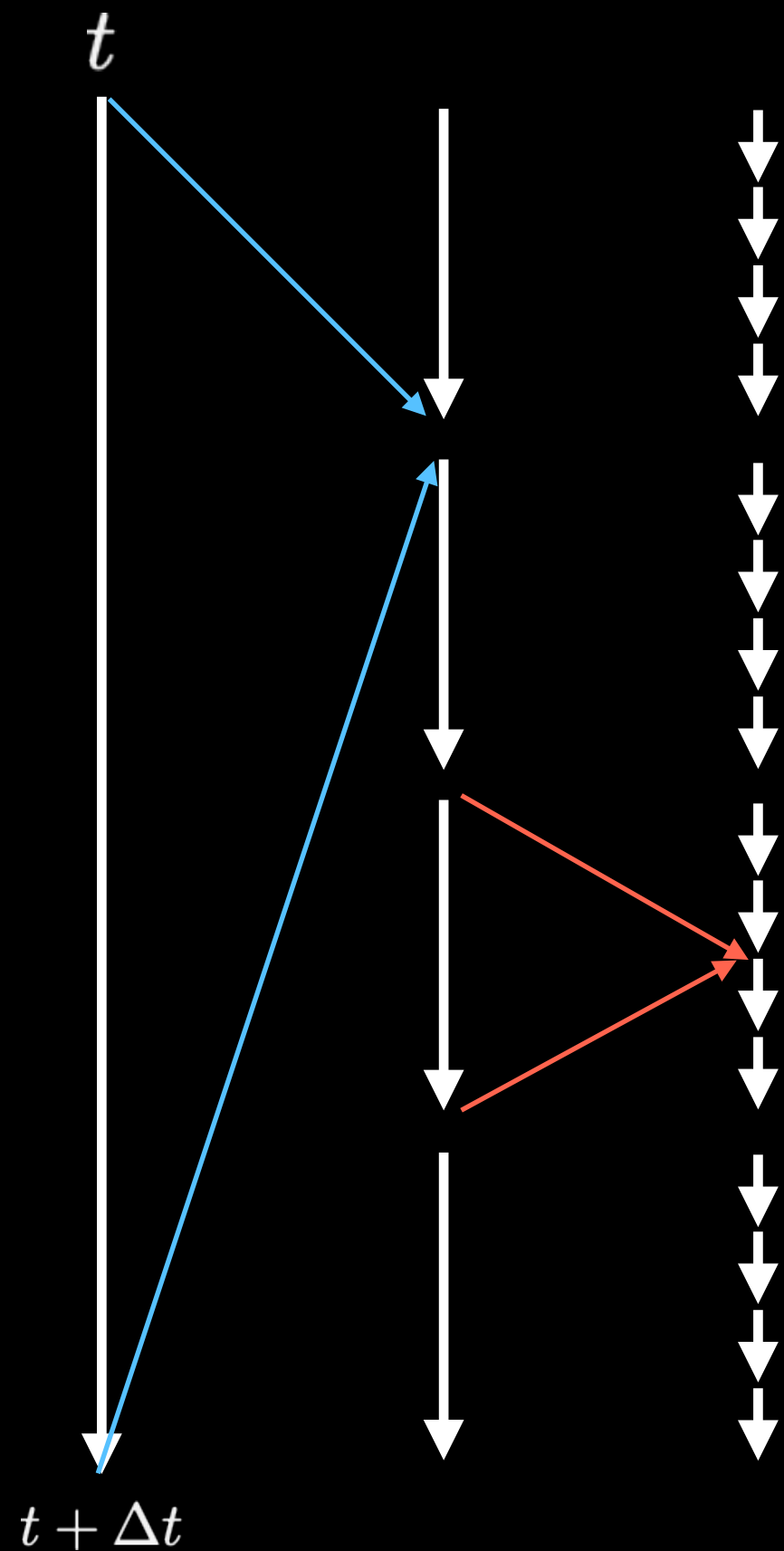
$$\mathbf{E}^{n+1} = -\mathbf{u}^{n+1} \times \mathbf{B}^{n+1} + \frac{\nabla \times \mathbf{B}^{n+1}}{N^{n+1}} - \frac{\nabla \cdot \mathbf{P}_e}{N^{n+1}} + \eta \nabla \times \mathbf{B}^{n+1} - \nu \nabla^2 \nabla \times \mathbf{B}^{n+1}$$



Third party AMR lib :
LLNL/SAMRAI

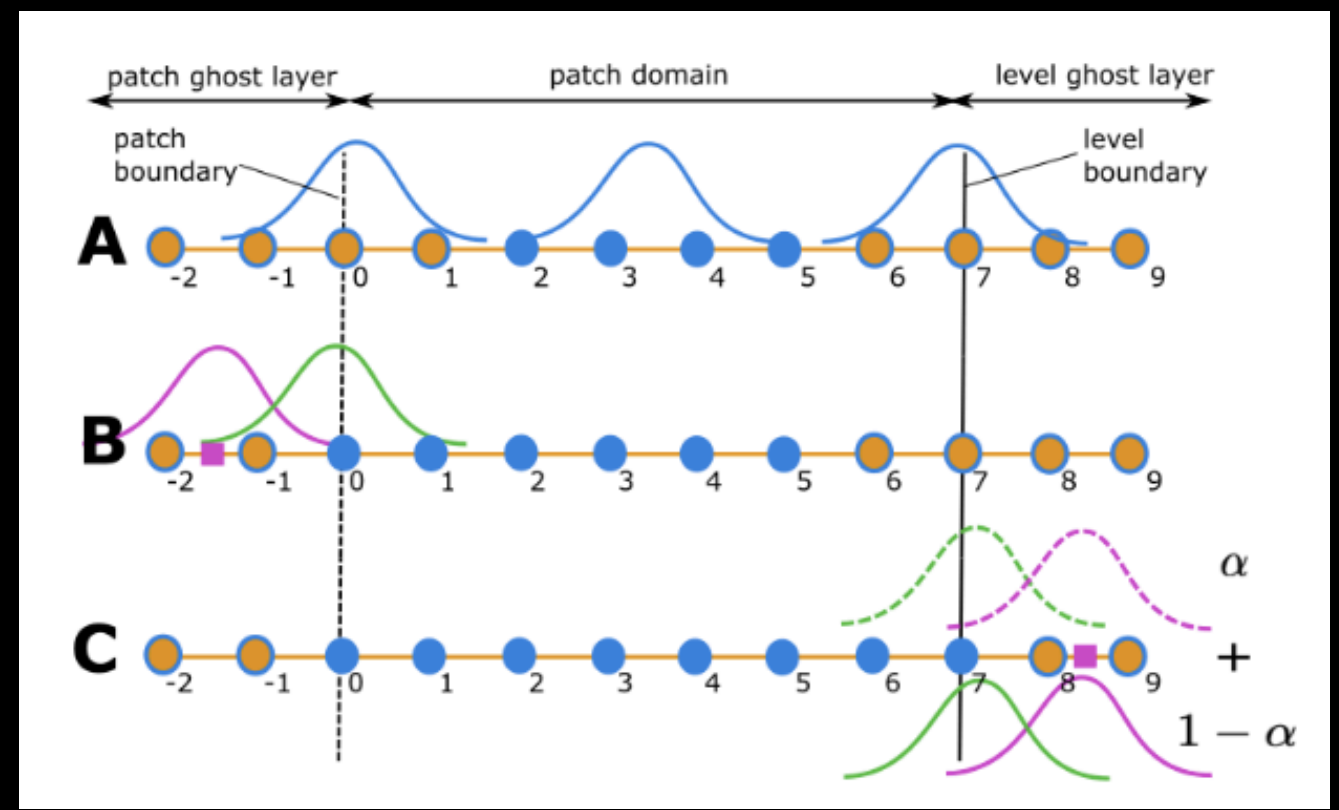
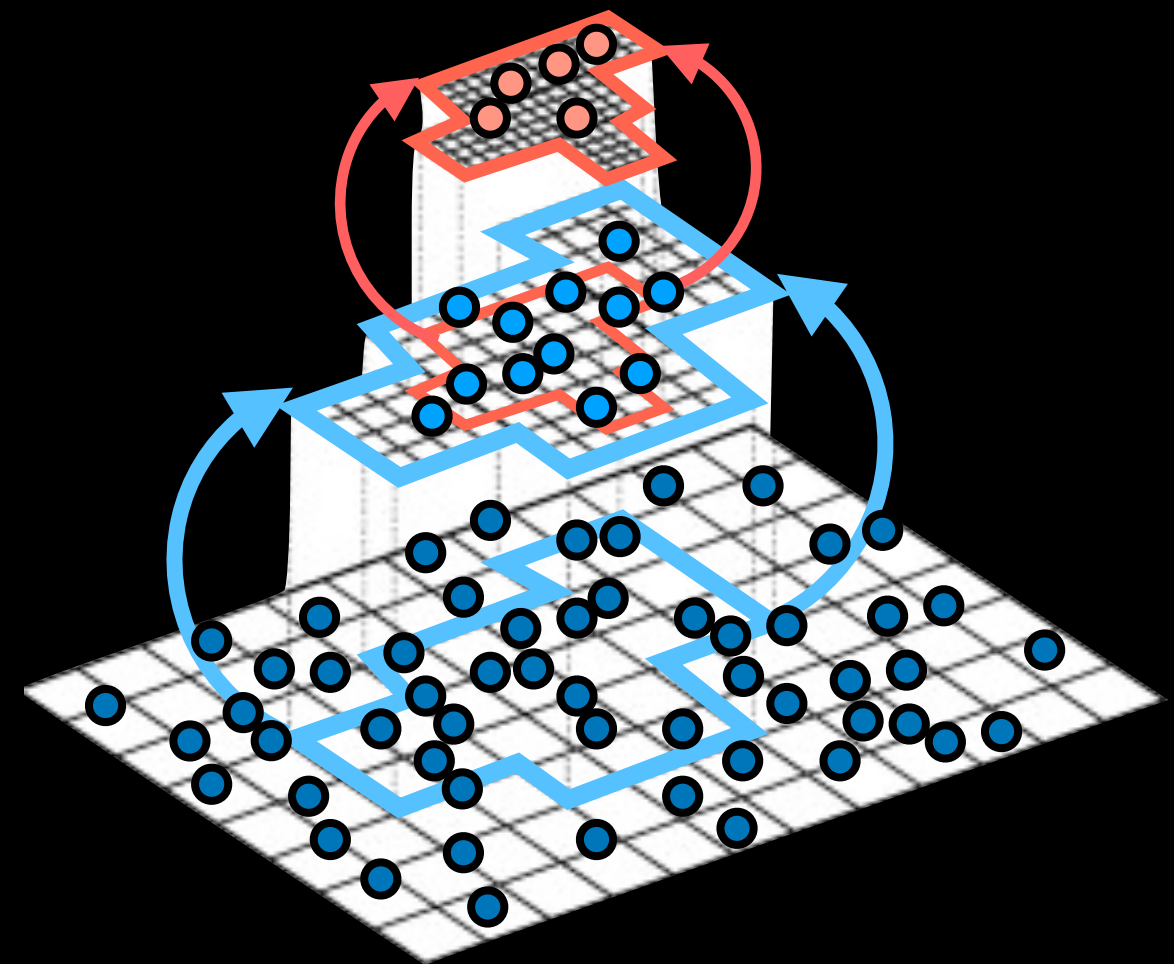
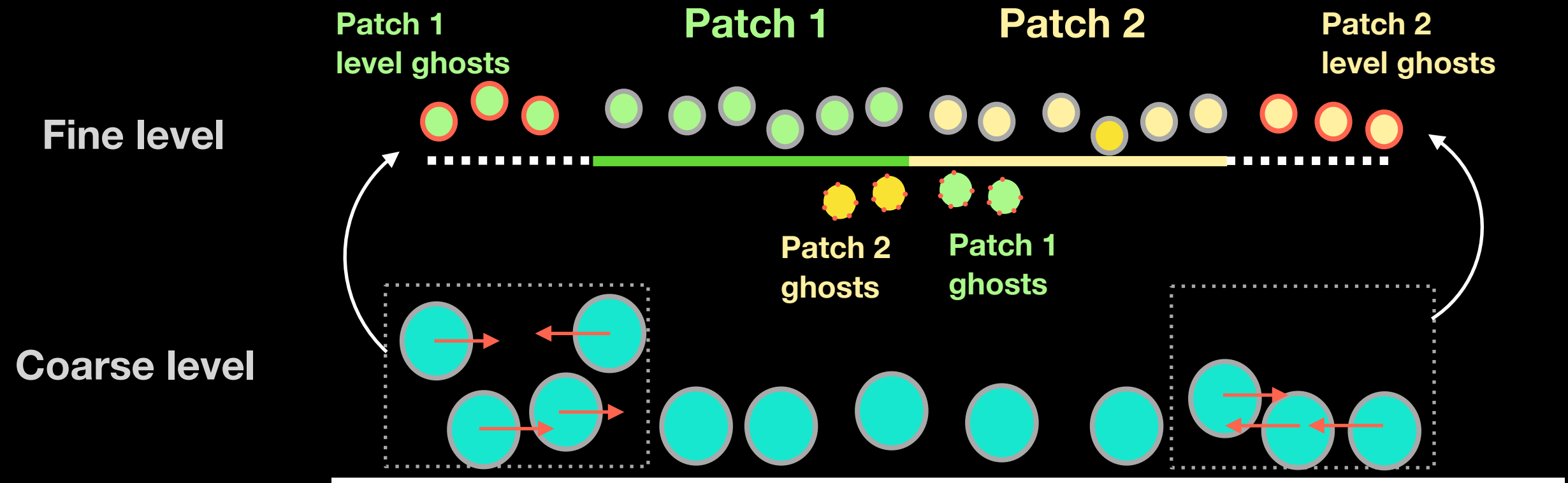
Coarse region are evolved 4x less often than next finer

COARSE SOLUTION IS TIME/SPACE INTERPOLATED ON FINE BOUNDARIES



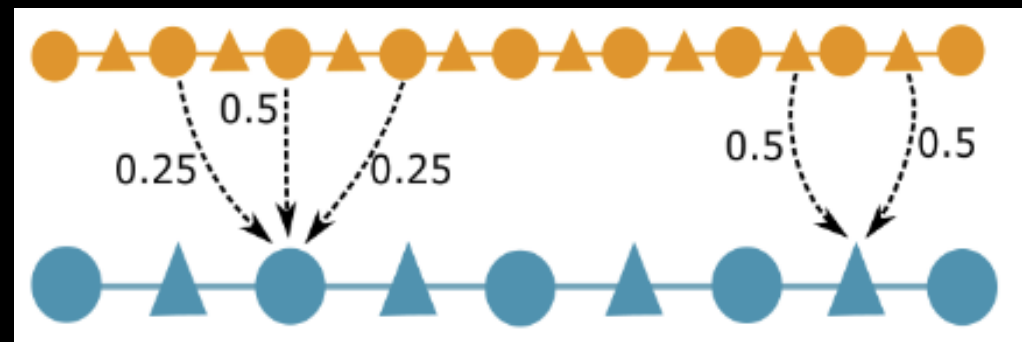
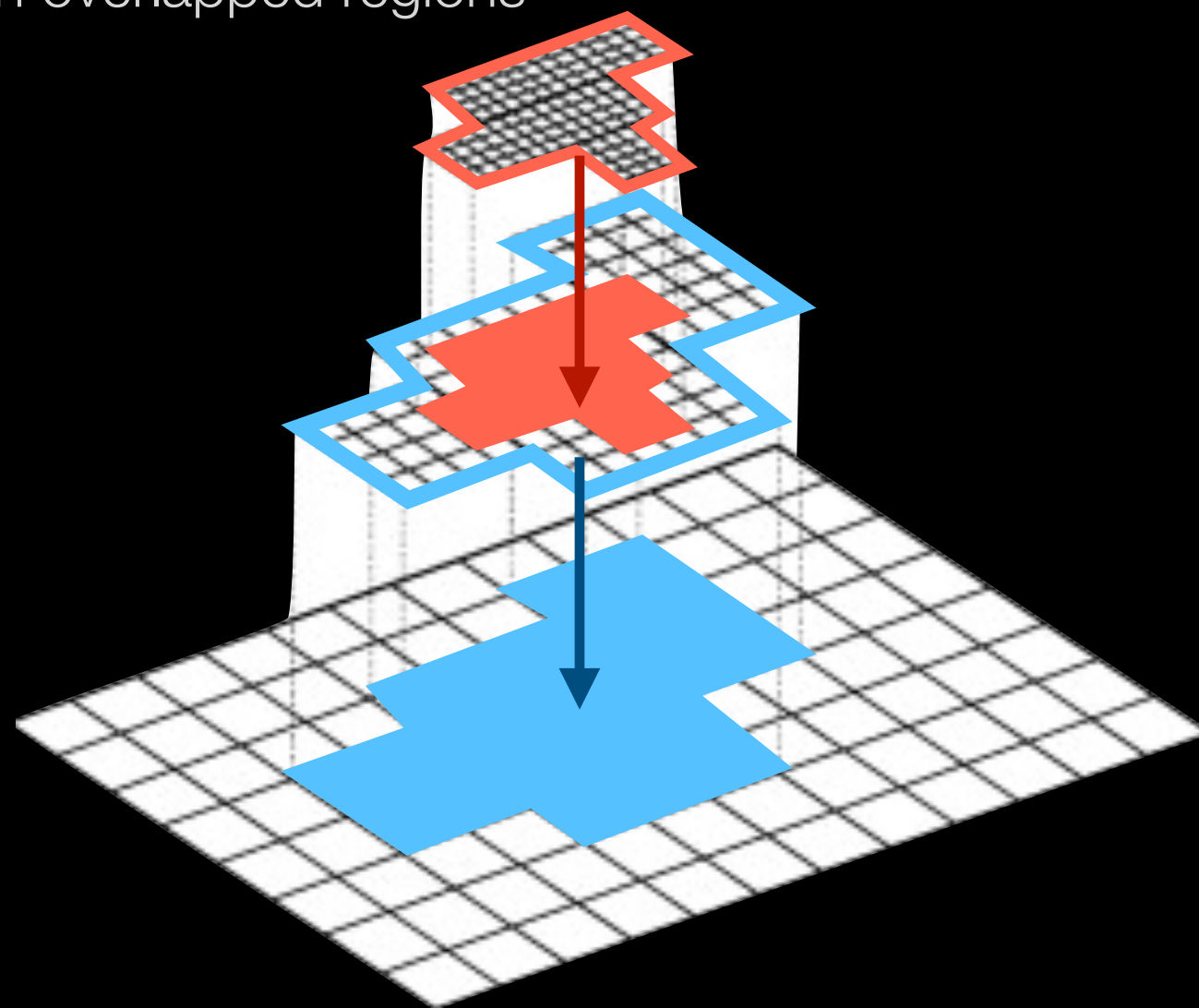
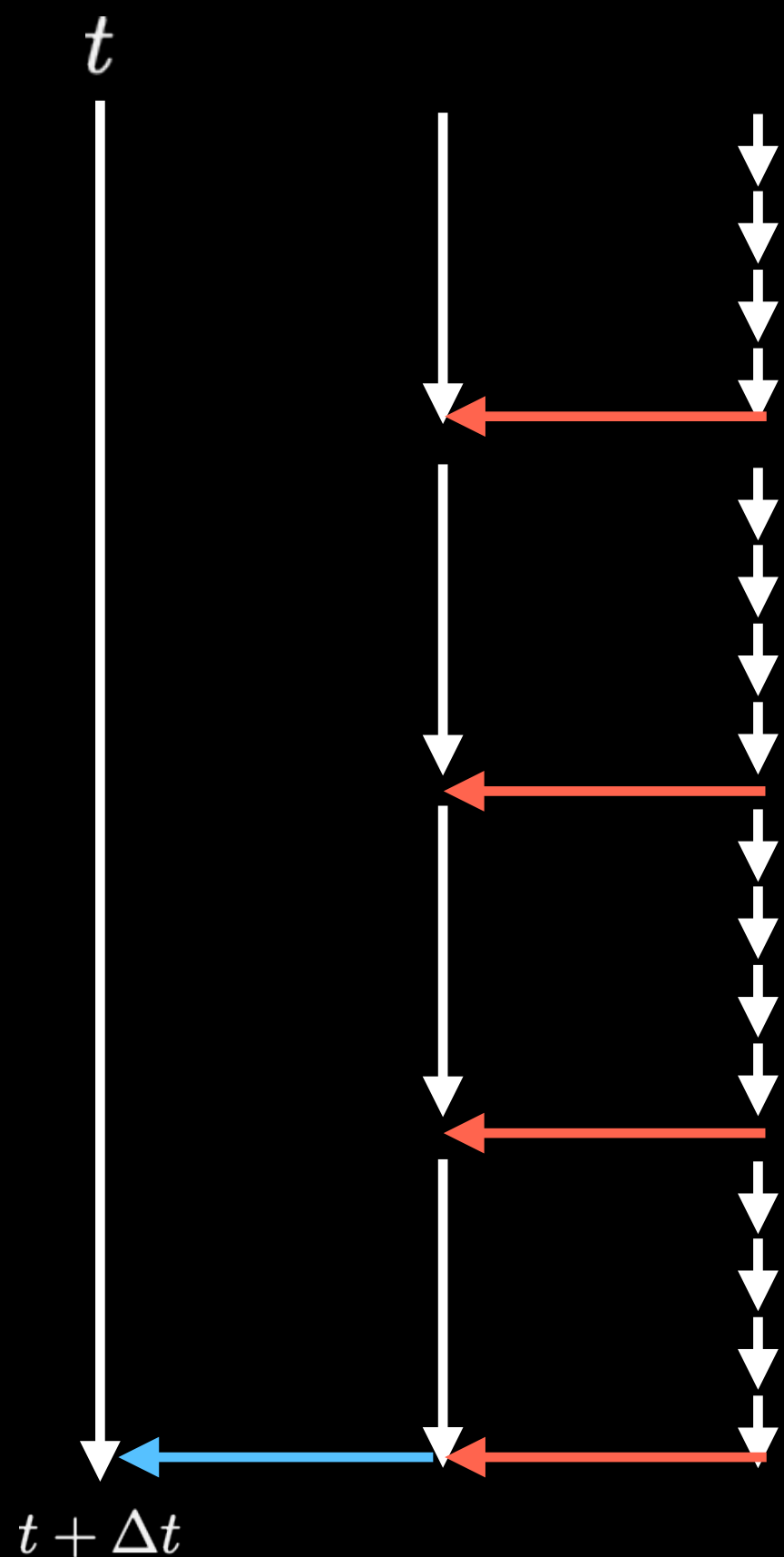
Coarse solution is linearly interpolated in time and then linearly interpolated in space on the refined grid.

COARSE SOLUTION IS TIME/SPACE INTERPOLATED ON FINE BOUNDARIES



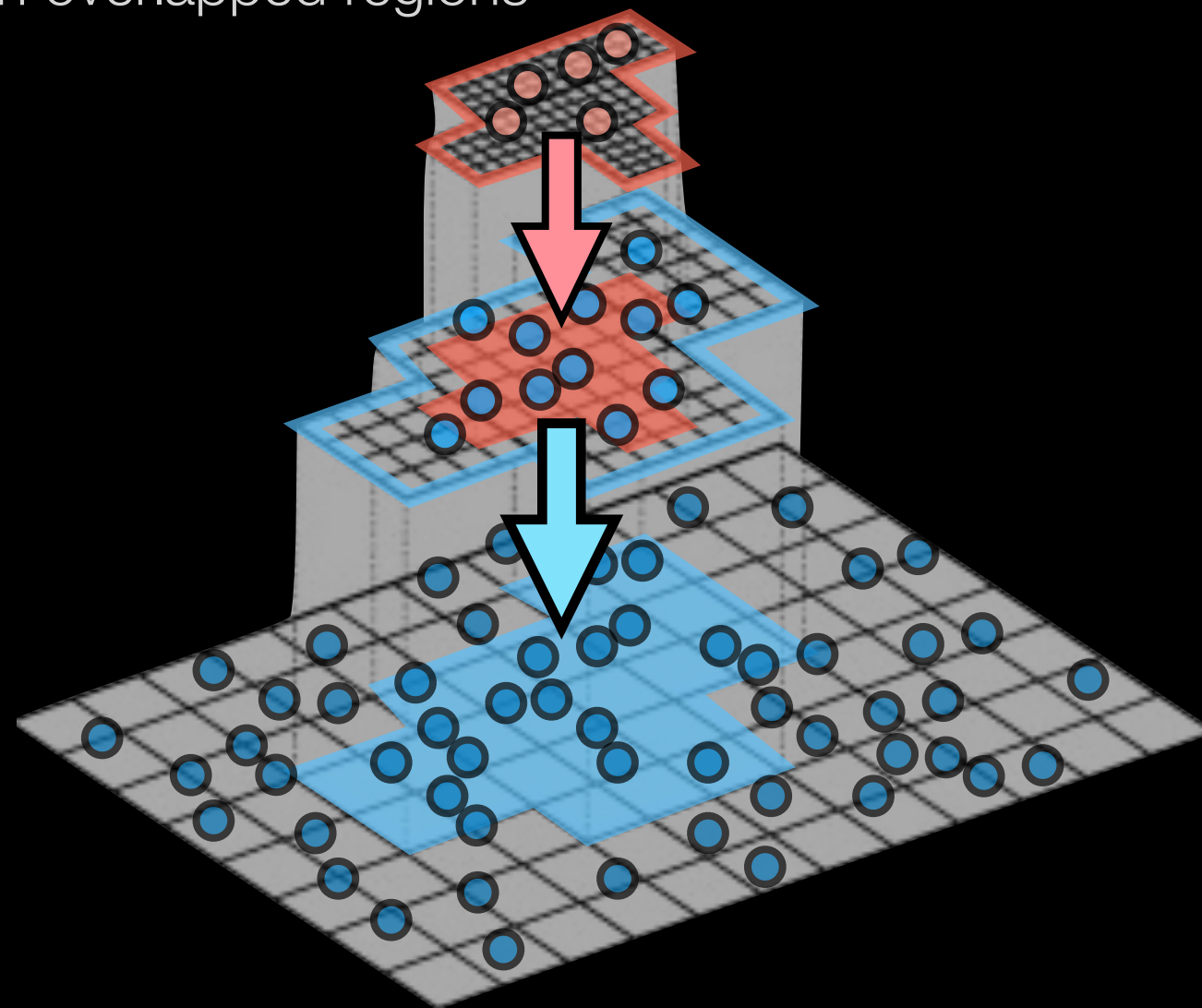
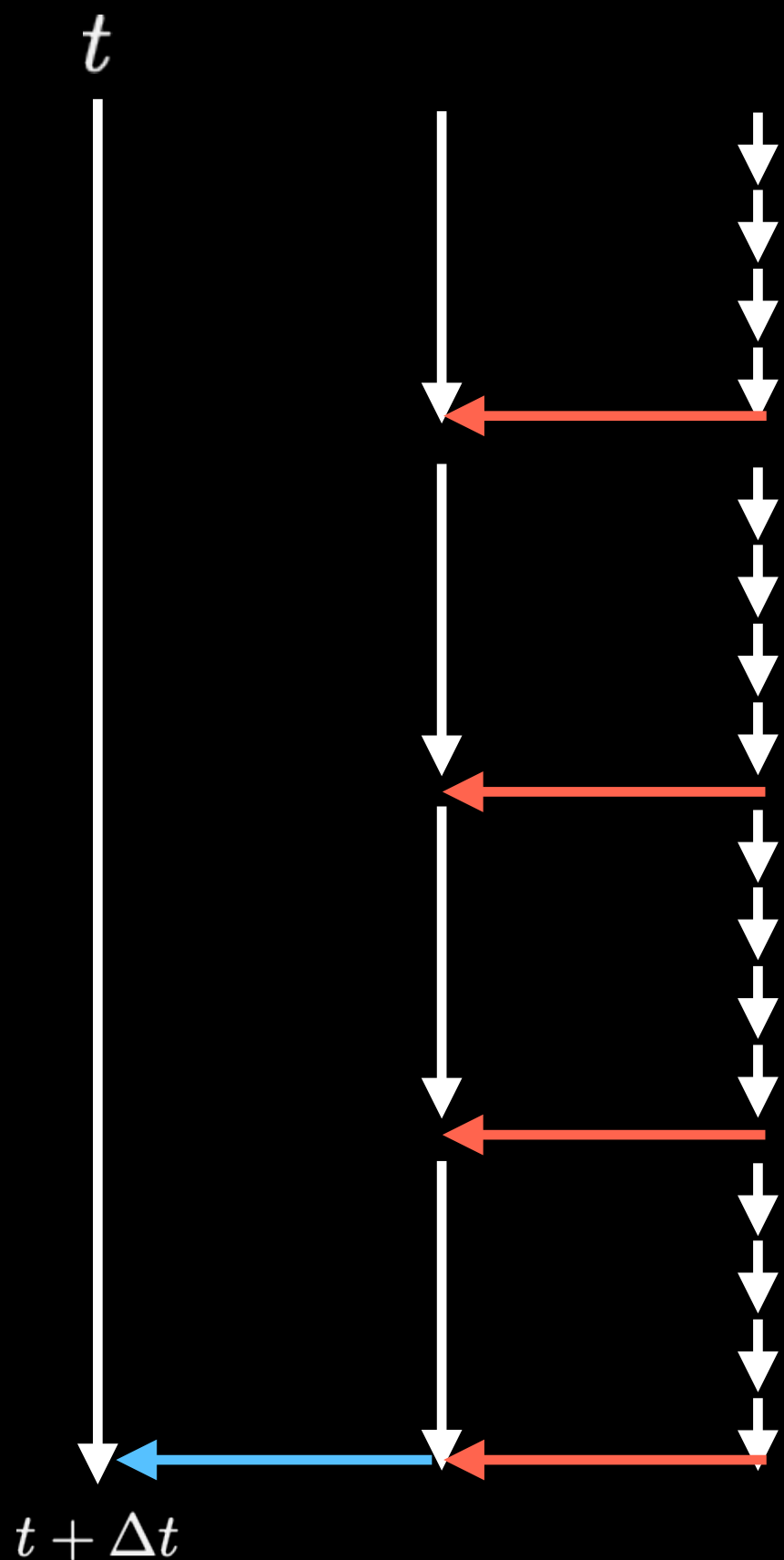
FINE SOLUTION ARE INTERPOLATED ON THE COARSE UNDERLYING PATCHES

Once a fine level has reached the next coarser time, the fine solution is coarsened and overwrites the coarse solution in overlapped regions



FINE SOLUTION ARE INTERPOLATED ON THE COARSE UNDERLYING PATCHES

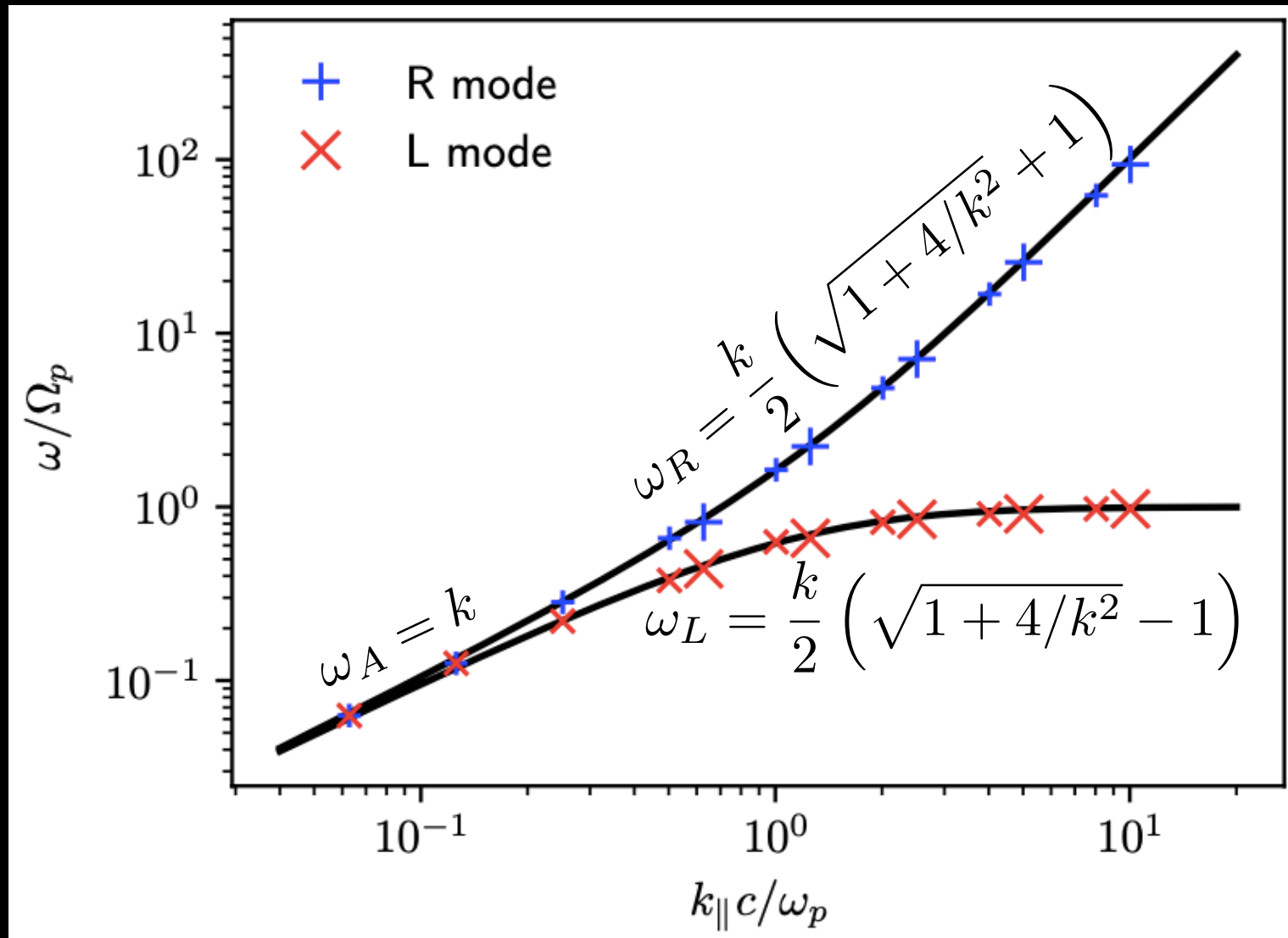
Once a fine level has reached the next coarser time, the fine solution is coarsened and overwrites the coarse solution in overlapped regions



Coarse particles evolve with electromagnetic fields constantly updated by the fine solution and stay synchronized

VALIDATION OF THE HYBRID SOLVER : WAVE DISPERSION

Dispersion diagram of parallel waves



whistler and ion cyclotron waves

$$\omega_{L,R} = \frac{k}{2} \left(\sqrt{1 + 4/k^2} \pm 1 \right)$$

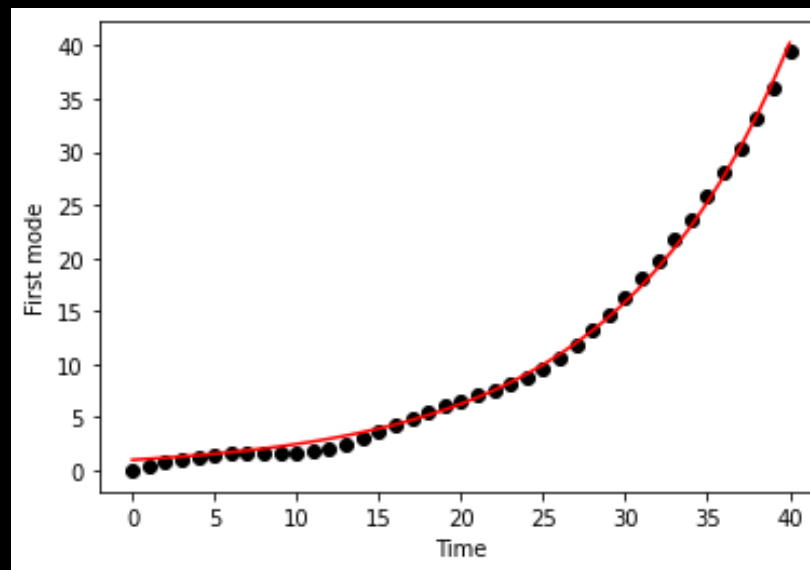
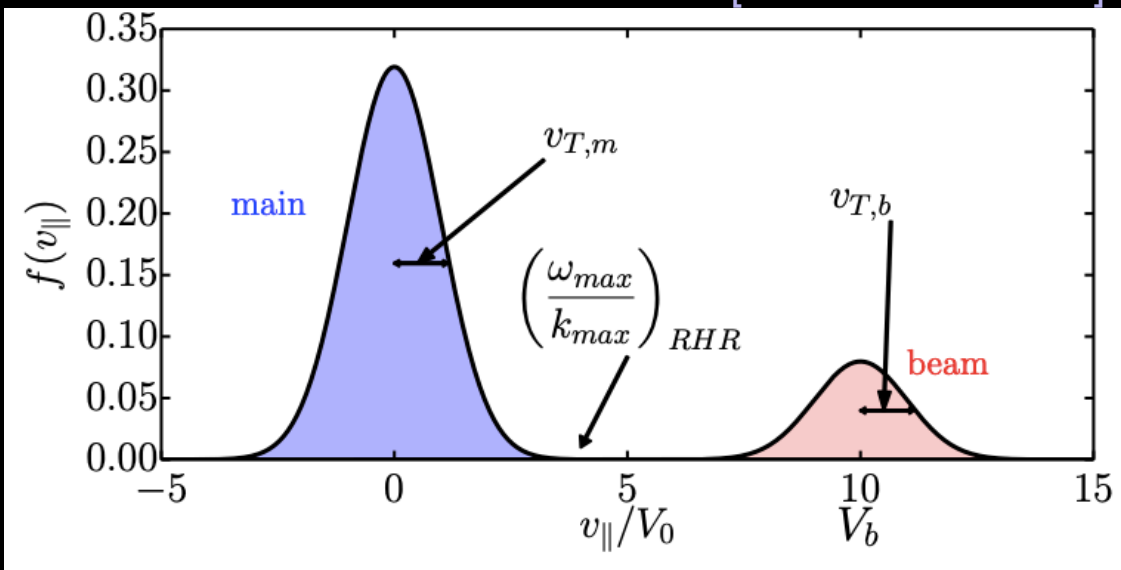
Alfvén waves

$$\omega_A = k$$

Small and large crosses represent 1D and 2D

1D VALIDATION : RIGHT HAND ION STREAMING INSTABILITY

[Nicolas et al. 2018]



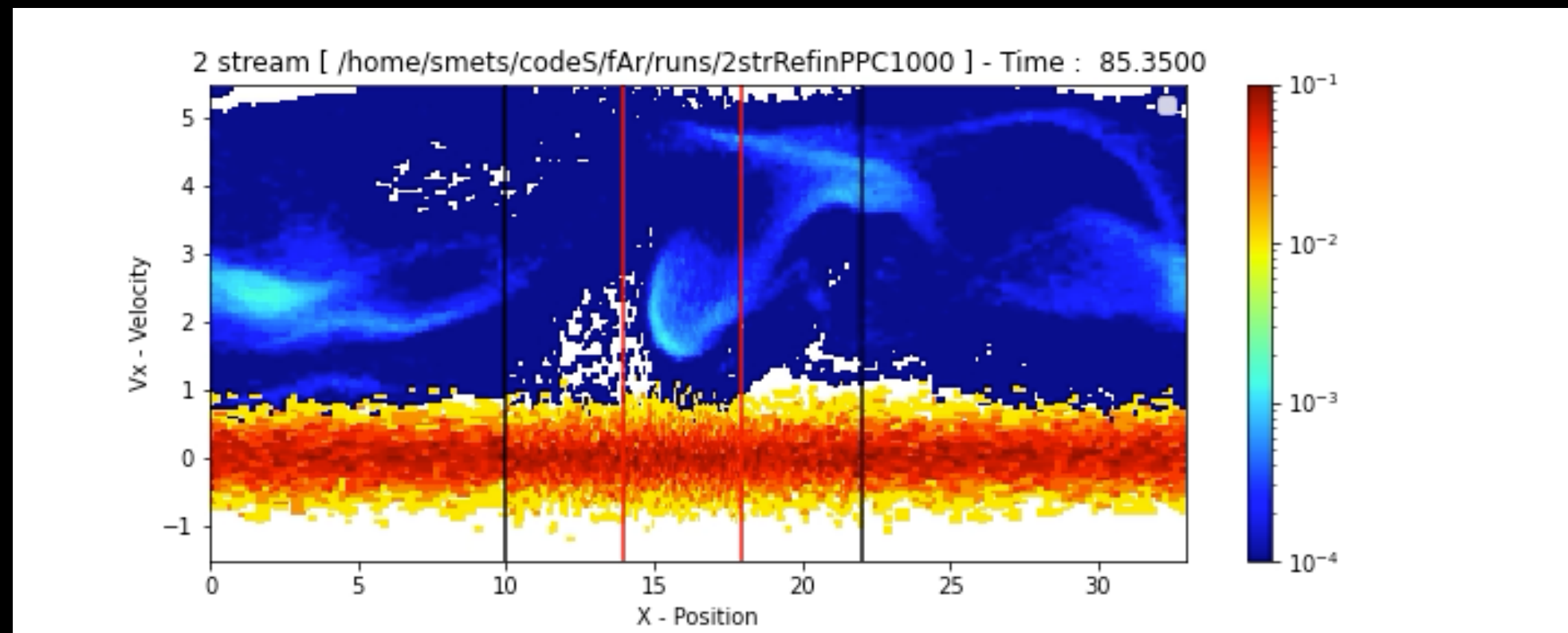
$$T_i = T_e = 0.1 \quad V_b = 5$$

$$n_m = 1 \quad n_b = 0.01$$

Linear theory $\gamma_{th} \approx 0.09$

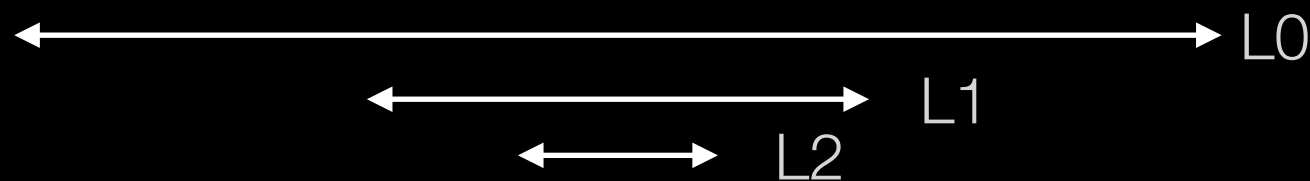
Model :

$$\gamma_{sim} \approx 0.0908 \pm 0.0053$$

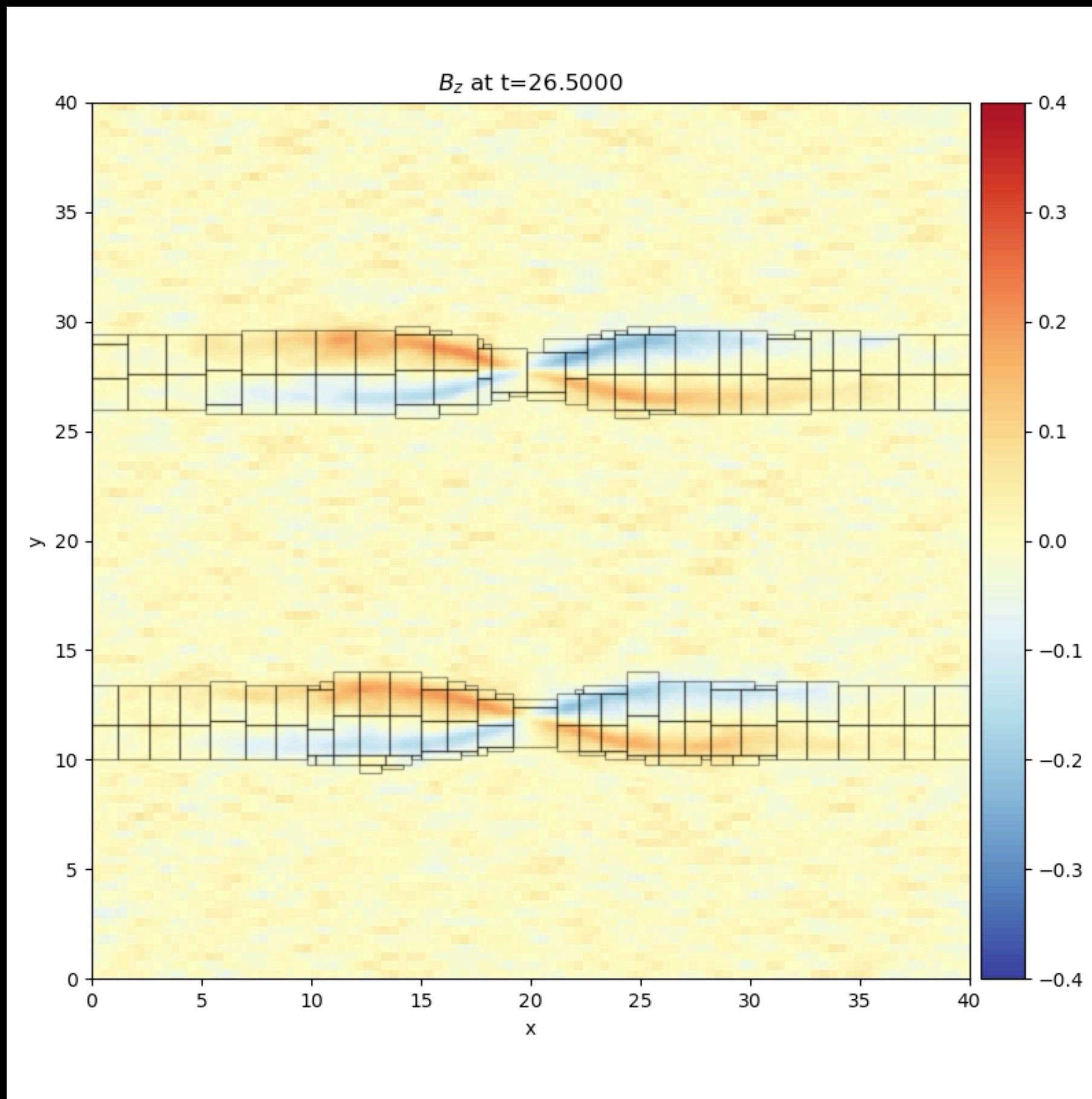


Phase space diagram

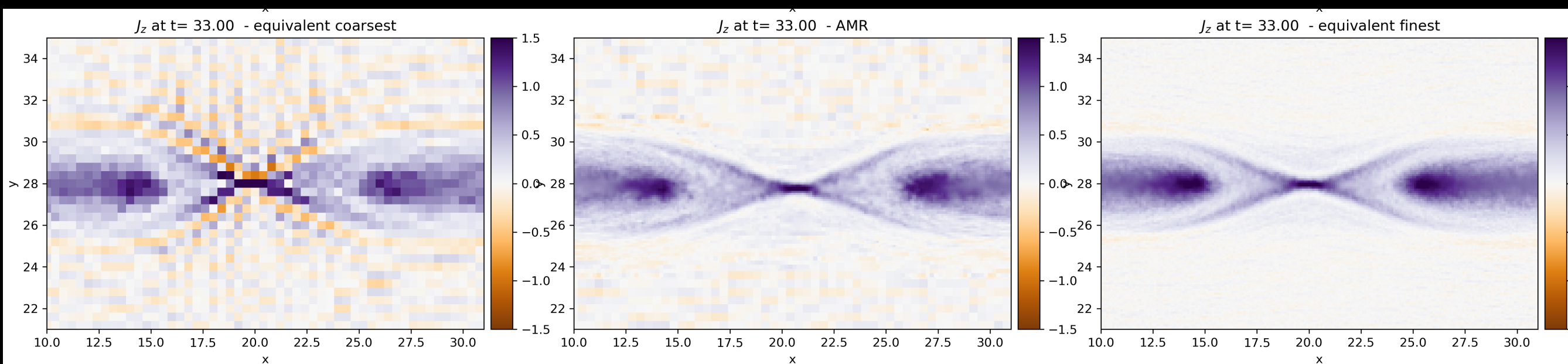
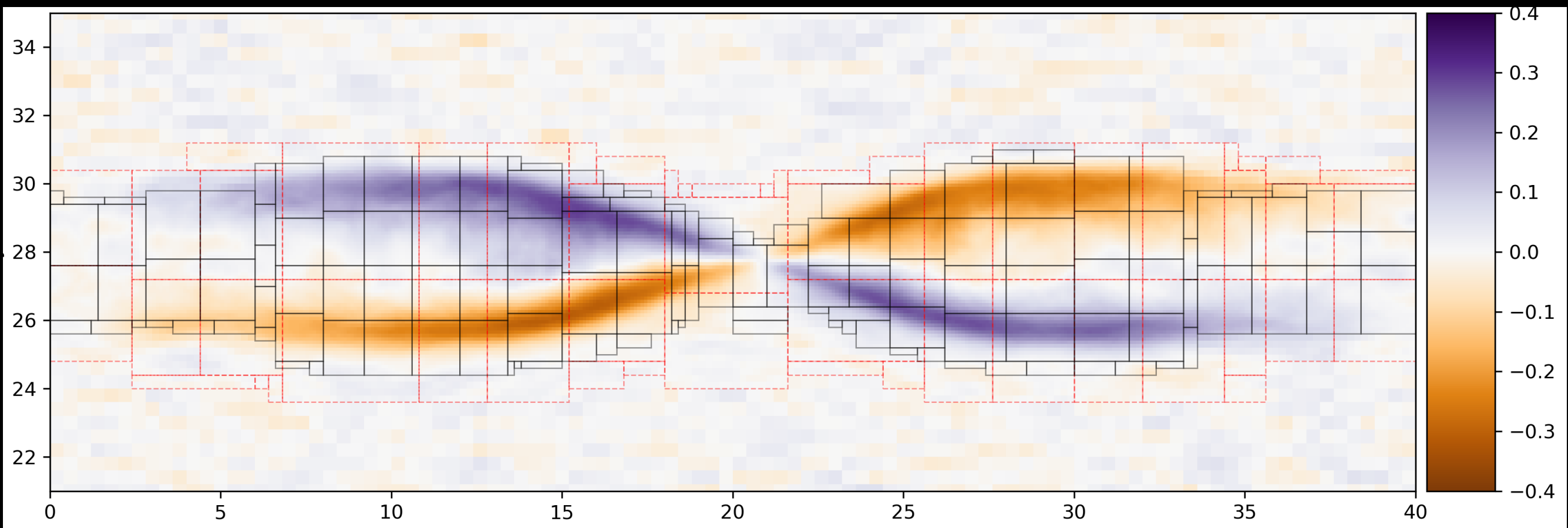
Structures cross level boundaries without any issue (ouf !)



2D MAGNETIC RECONNECTION



2D MAGNETIC RECONNECTION



[Aunai et al. in prep]



100% open source



High perf.
Abstractions



Ergonomy



Tested

- Aim at **high performance** while maintaining a **super user friendly** (python) interface
- Target users: simulation **experts and non experts** (observers, students...), classes
- Extensive unit (>1000) and functional tests at each merge and nightly builds to get a **robust code** that **can be upgraded and maintained**